

Neutrino Oscillation Experiments at Reactors and Accelerators

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IV^{èmes} Rencontres du Vietnam, Hanoi, July 2000

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CHOOZ and **PALO VERDE** long baseline experiments

Search for ν_e disappearance at reactors.

Large mixing angle, small Δm^2 ($>10^{-3}$ eV²)

KARMEN2 and **LSND** short baseline experiments

Search for $\nu_\mu - \nu_e$ oscillation in low energy neutrino beams.

High sensitivity, small mixing angle, large Δm^2 (few eV²)

CHORUS and **NOMAD** short baseline experiments

Search for $\nu_\mu - \nu_\tau$ oscillation in high energy neutrino beams.

High sensitivity, small mixing angle, large Δm^2 (few tens eV²)

To-morrow Byron Lundberg will give a seminar at FermiLab on

"Results from DONUT: First Direct Evidence of the Tau Neutrino"

This was expected and did not happen explicitly in Sudbury at
Neutrino 2000

Neutrino Oscillation Experiments at Nuclear Reactors

Long base line experiments

at nuclear power plants of Chooz and Palo Verde

Motivation:

$\bar{\nu}_e$ disappearance

in $(\Delta m^2, \sin^2 2\theta)$ parameter space indicated by

ν_μ disappearance in atmospheric experiments $\nu_\mu \rightarrow \nu_e$?

Neutrino Oscillation at Reactors: pros and cons

- $E_\nu \approx \text{few MeV} \Rightarrow$ Access to low Δm^2 at medium L $\Delta m^2 \approx \frac{E \approx 3 \text{ MeV}}{L \approx 1000 \text{ m}} \approx 0.003 \text{ eV}^2$

\Rightarrow Below μ, τ thresholds: only disappearance $\bar{\nu}_e \rightarrow \bar{\nu}_x$

- High flux, but small σ

- 4π source \Rightarrow detector mass $\div L^2$

$$P = 1 \text{ GW}_{elec} \Rightarrow 3.3 \text{ GW}_{therm}$$

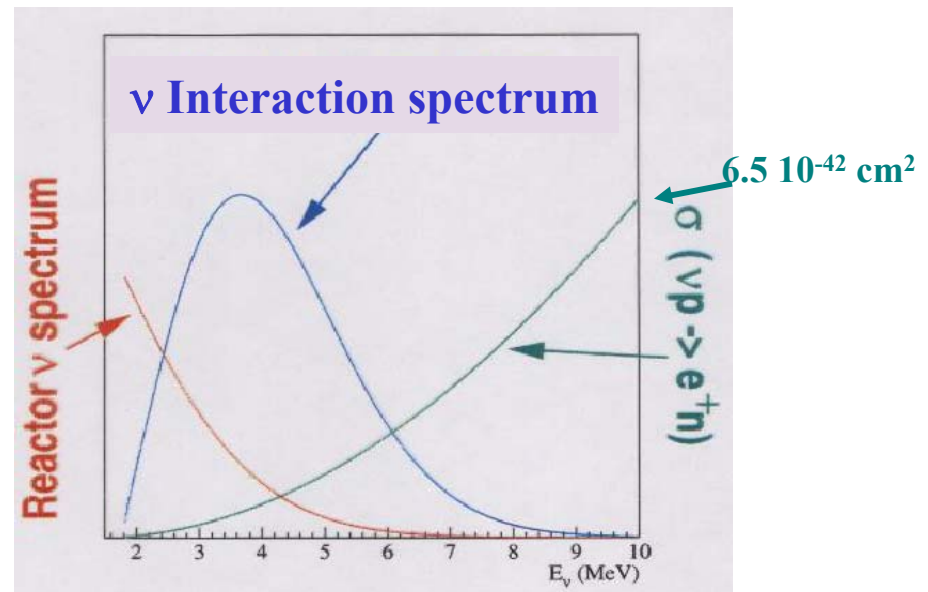
$$\text{Fission rate} \approx \frac{3.3 \text{ GW}}{200 \text{ MeV}} \approx 10^{20} \text{ s}^{-1} \Rightarrow 6 \cdot 10^{20} \bar{\nu}_e \text{ s}^{-1}$$

- Disappearance \Rightarrow good knowledge of absolute ν flux and e^+ energy spectrum
 \Rightarrow or multi-L experiment (≥ 2 detectors or reactors) \Rightarrow
no sensitivity at high Δm^2
 (not serious problem with $\Delta m^2 \leq 0.01 \text{ eV}^2$)

- Cheap and well known ν source
 Calculated and measured ν flux
 and energy spectrum at $L=0$ known
 to $\sim 2\%$ (Bugey 1995)

$$E_{\bar{\nu}_e}^{thresh} = 1.8 \text{ MeV}$$

$$\langle E_{e^+} \rangle \approx 3 \text{ MeV}$$



Detection of neutrinos from nuclear reactors

1953 : F.Reines and C.L. Cowans discover the neutrino
at Savannah River nuclear power plant

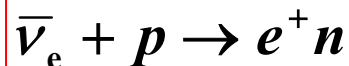
Detectors

- vessel filled with liquid scintillator doped with neutronphage
- shielding (bunker, underground)
+ active veto: cosmic rays, reactor n,
natural radioactivity

Signal

$$E_{\bar{\nu}_e}^{thresh} = 1.8 \text{ MeV}$$
$$\langle E_{e^+} \rangle \approx 3 \text{ MeV}$$

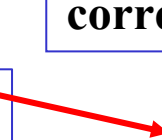
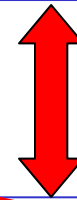
nuclear capture $\rightarrow \gamma$'s ($\sum E_\gamma$ known)



scintillation γ 's
 $e^+ - e^-$ annihilation : 2 γ of 0.511 keV

Space and
delayed time
correlation

Cerenkov light



The Long Base Line CHOOZ Experiment

M. Apollonio^c, A. Baldini^b, C. Bemporad^b, E. Caffau^c, F. Cei^b, Y. Déclais^{e,1},
H. de Kerret^f, B. Dieterle^h, A. Etenko^d, J. George^h, G. Giannini^c, M. Grassi^b,
Y. Kozlov^d, W. Kropp^g, D. Kryn^f, M. Laiman^e, C.E. Lane^a, B. Lefèvre^f,
I. Machulin^d, A. Martemyanov^d, V. Martemyanov^d, L. Mikaelyan^d, D. Nicolò^b,
M. Obolensky^f, R. Pazzi^b, G. Pieri^b, L. Price^g, S. Riley^g, R. Reeder^h,
A. Sabelnikov^d, G. Santin^c, M. Skorokhvatov^d, H. Sobel^g, J. Steele^a, R. Steinberg^a,
S. Sukhotin^d, S. Tomshaw^a, D. Veron^f, and V. Vyrodov^f

^a*Drexel University*

^b*INFN and University of Pisa*

^c*INFN and University of Trieste*

^d*Kurchatov Institute*

^e*LAPP-IN2P3-CNRS Annecy*

^f*PCC-IN2P3-CNRS Collège de France*

^g*University of California, Irvine*

^h*University of New Mexico, Albuquerque*

¹*Present address: IPNL-IN2P3-CNRS Lyon*

Phys. Let. B466 (1999) 415

CHOOZ detector

-1 detector - 2 reactors (8.5GW) : $L=998, 1114\text{m}$
 $\Delta L=116.7\text{m}$

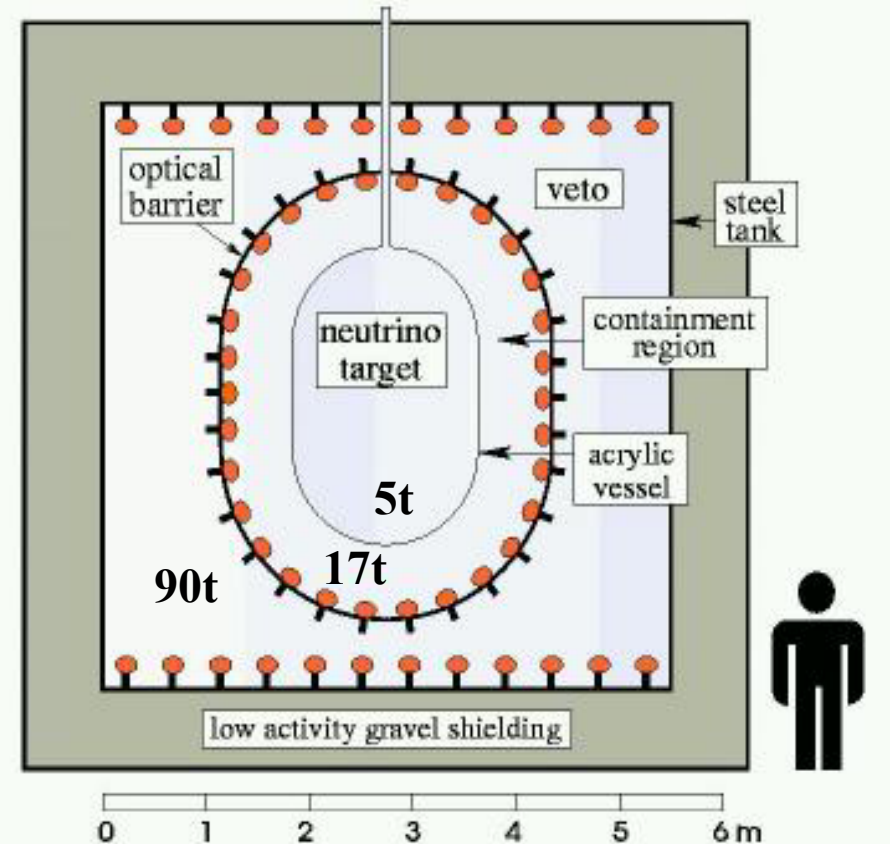
-rock overburden: 300 m water equivalent
 $0.4 \text{ cosmic } \mu \text{ m}^{-2} \text{ s}^{-1}$

-5 tons Gd-doped liquid scintillator (0.09%)

$$\sum E_{\gamma} = 8 \text{ MeV}$$

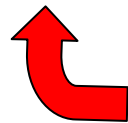
-17 tons liquid scintillator : contain γ from n
 PMT radioactivity shield

-90 tons active cosmic-ray muon veto



Event rates: full power: 24.7 ± 0.7 events/day
reactors off: 1.2 events/day

Data taking: April 1997 - July 1998		
Reactor 1 ON	2058.0 h	8295 GWh
Reactor 2 ON	1187.8	4136
Reactors 1 & 2 ON	1543.1	8841
Reactors OFF	3420.4	



Background estimates

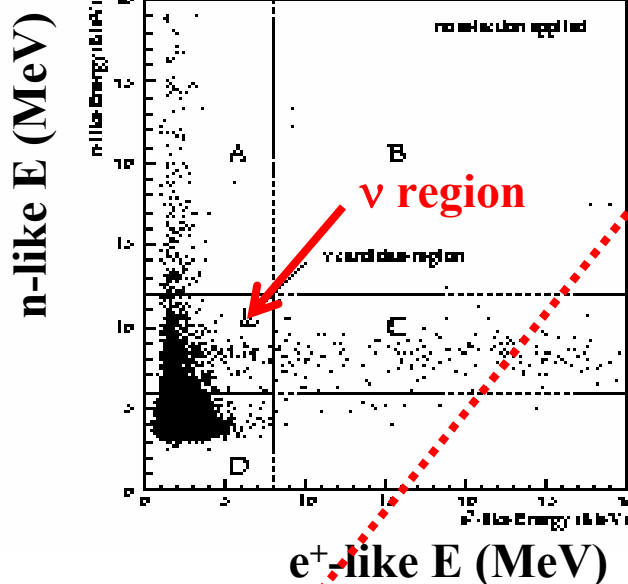
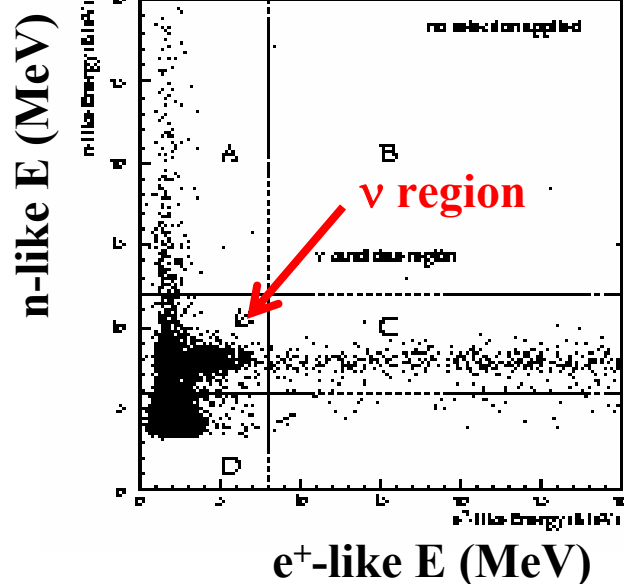
Response calibration: γ , n and γ -n radioactive sources (^{60}Co , ^{252}Cf , Am/Be)

E_n^{abs} time dependence monitoring ($\sum E_\gamma = 8 \text{ MeV}$) with n from cosmic : $\sigma_E = 0.5 \text{ MeV}$

Reactor ON

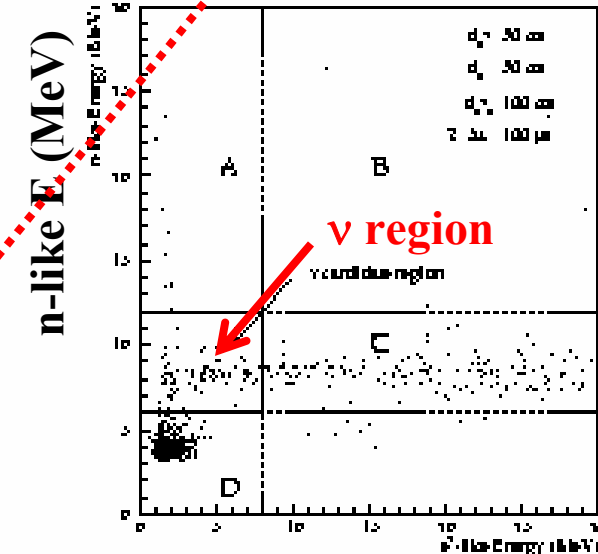
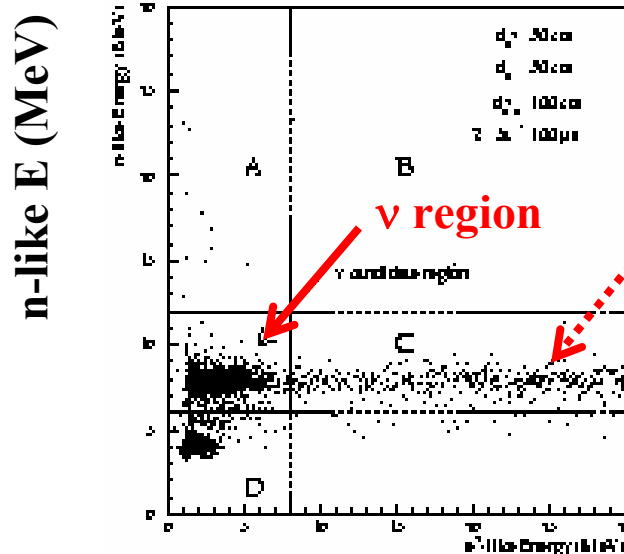
Reactor OFF

No event selection



Main background

fast spallation n in rock
+
p from n scattering (e⁺ like)
+
n capture



ν selection

@ > 30 cm from wall,
n - e⁺ distance < 100 cm
n - e⁺ delay in (2-100) μ s
E(e⁺-like) in (1.3 - 8) MeV
E(n-like) in (6-12) MeV

**2991 candidates
(287 reactors off)**

Efficiency: 69.8%

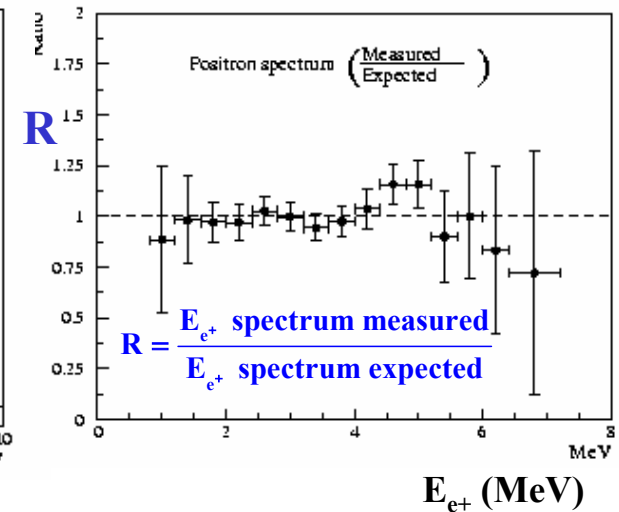
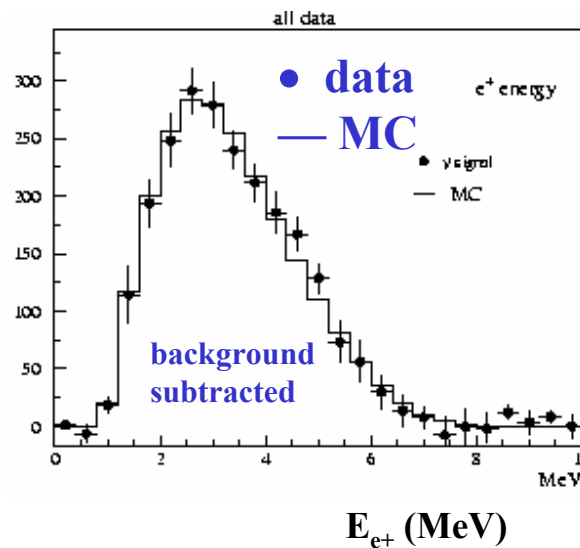
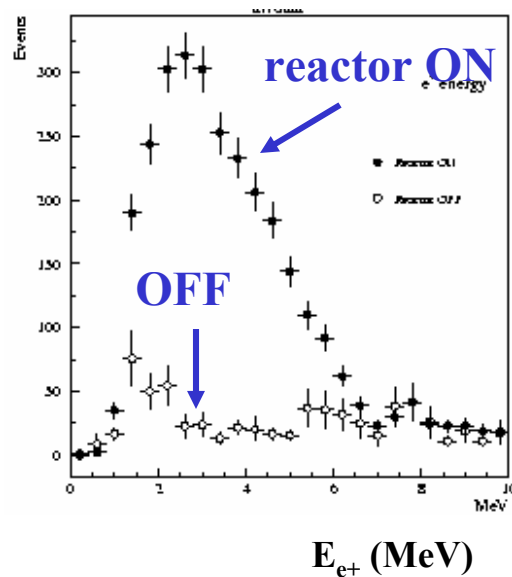
$\bar{\nu}_e$ flux known to 1.4%

- daily evolution of core isotopic evolution
- instantaneous fission rate from thermal power
- ν yield from measured β spectra of main isotopes

+

E_{e^+} spectrum

- inverse β -decay cross-section
- simulation of detector response



$$R = 1.010 \pm 0.028 \text{ (stat)} \pm 0.027 \text{ (syst)}$$



No oscillation signal

Analysis Methods

A - Compare unfolded E_{e^+} absolute spectra of both reactors to expectation

Systematic uncertainty on absolute normalisation: $\sim 2\%$

Two “independent” measurements

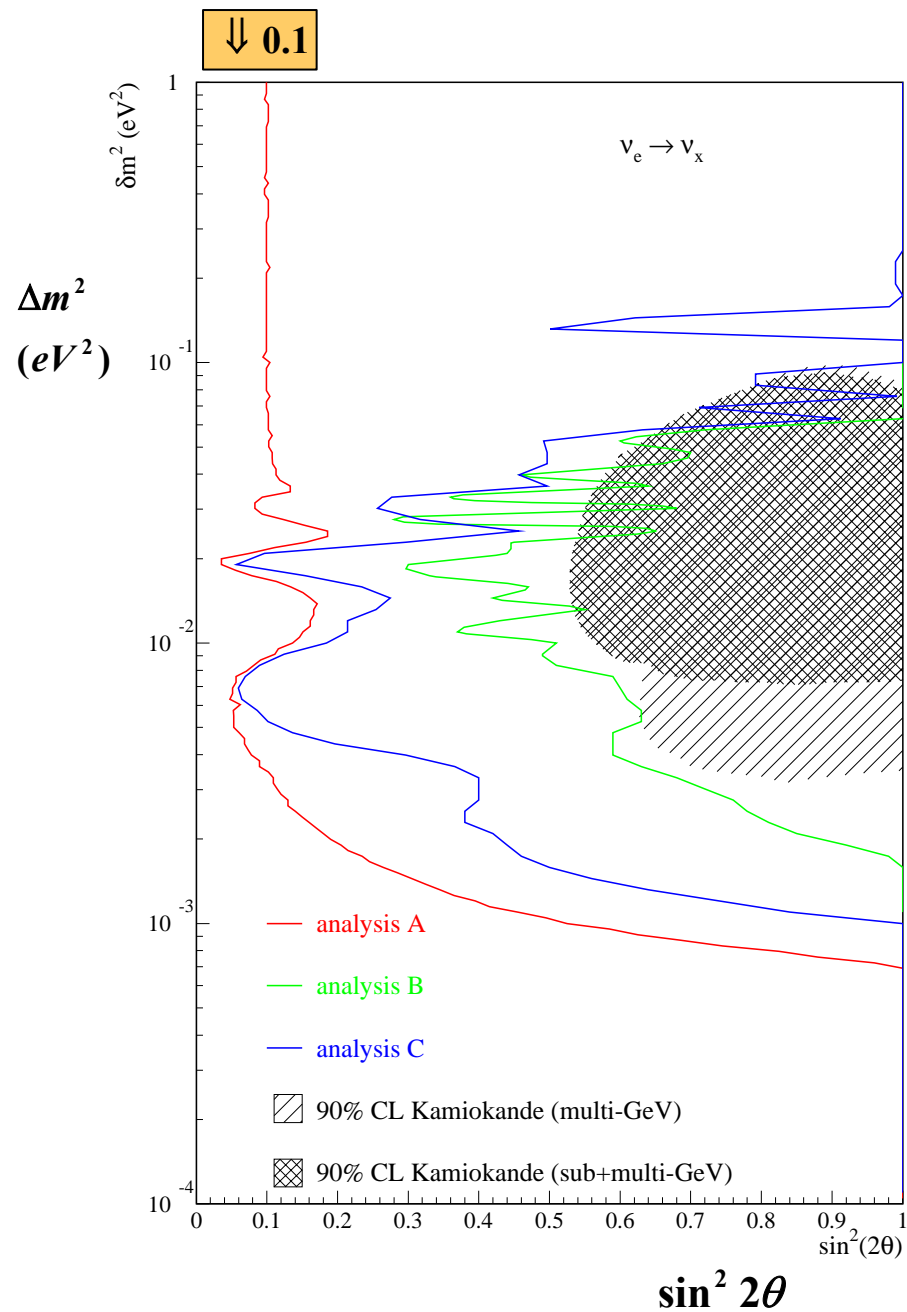
B - Ratio of spectra

Most systematic cancel

No sensitivity at large Δm^2

C - Compare unfolded E_{e^+} spectra shapes of both reactors to expectation

Intermediate sensitivity



Chooz exclusion plot

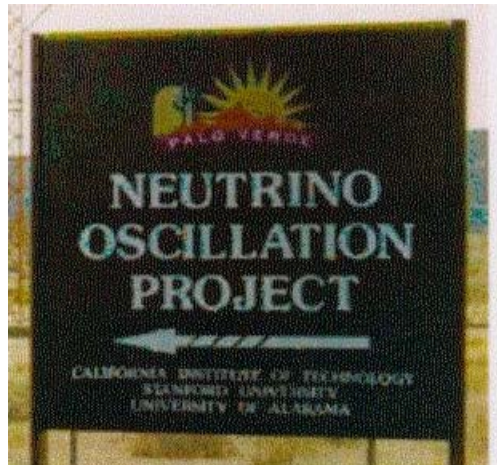
A — absolute spectra

B — spectra ratio

C — spectra shape

Kamiokande 90%

The Long Base Line Palo Verde Experiment



J. Busenitz, J. Kornis, K. McKinny, J. Wolf
University of Alabama

D. Lawrence, B. Ritchie
Arizona State University

F. Boehm, B. Cook, H. Henrikson, V. Novikov,
A. Piepke, P. Vogel, K.B. Lee
Caltech

G. Gratta, L. Miller, D. Tracy, Y-F. Wang
Stanford University

G.Gratta Neutrino 2000
F.Boehm et al. hep-ex/000322

-1 detector - 3 reactors (**11.6 GW**) : L= 750, 890m
 $\Delta L = 110\text{m}$

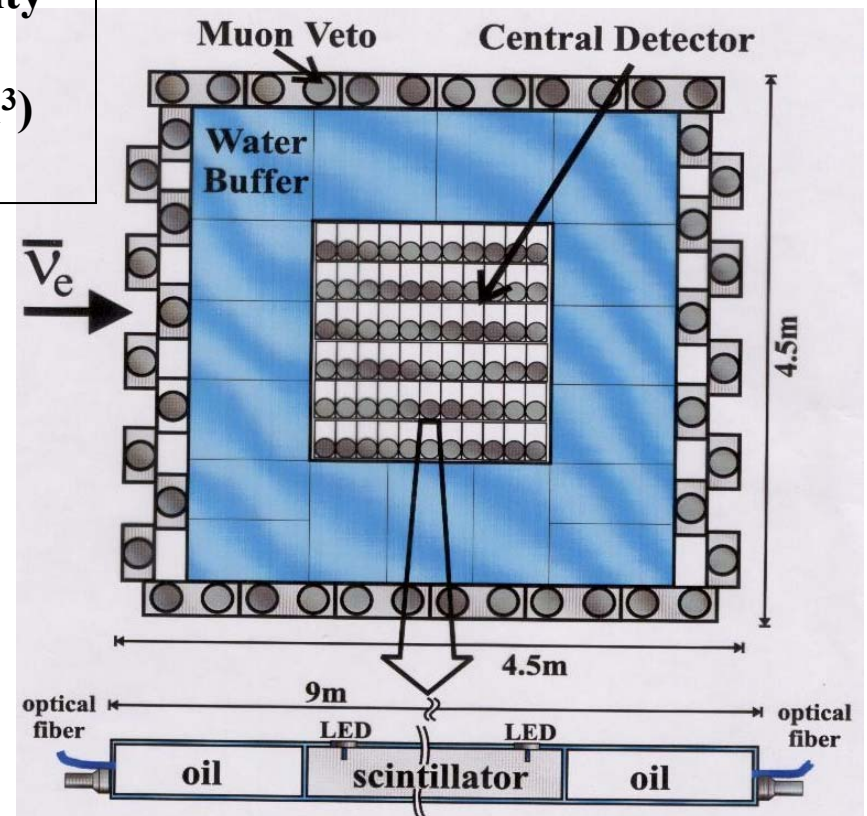
-rock overburden: **32 m water equivalent**
22 cosmic μ $\text{m}^{-2} \text{s}^{-1}$

-**11.3 tons** Gd-doped liquid scintillation (0.1%)
 $\sum E_{\gamma} = 8 \text{ MeV}$

- oil and 105 tons water buffer: γ and n shield
 shield PM radioactivity

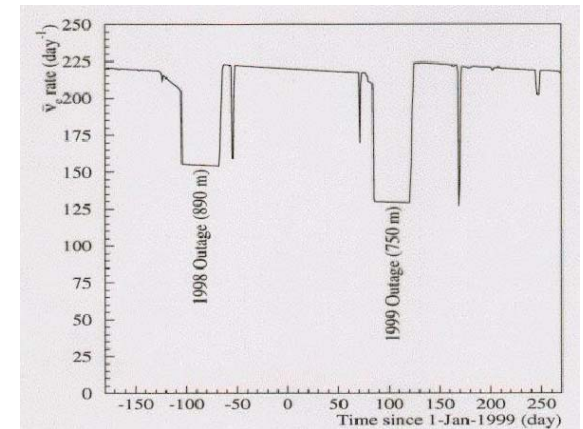
-**optically segmented detector** ($900 \times 12.7 \times 25.4 \text{ cm}^3$)
 \Rightarrow background suppression

Palo Verde detector



Difficulty : No period with all reactors off to measure simply the reactors off background.

Analysis based on the knowledge of the flux form
the known reactors power \Rightarrow True expected event number
compared to
Observed number of candidates
corrected for detector efficiencies (MC)



Unknowns :

-Background

-Overall normalisation
within systematic
uncertainty



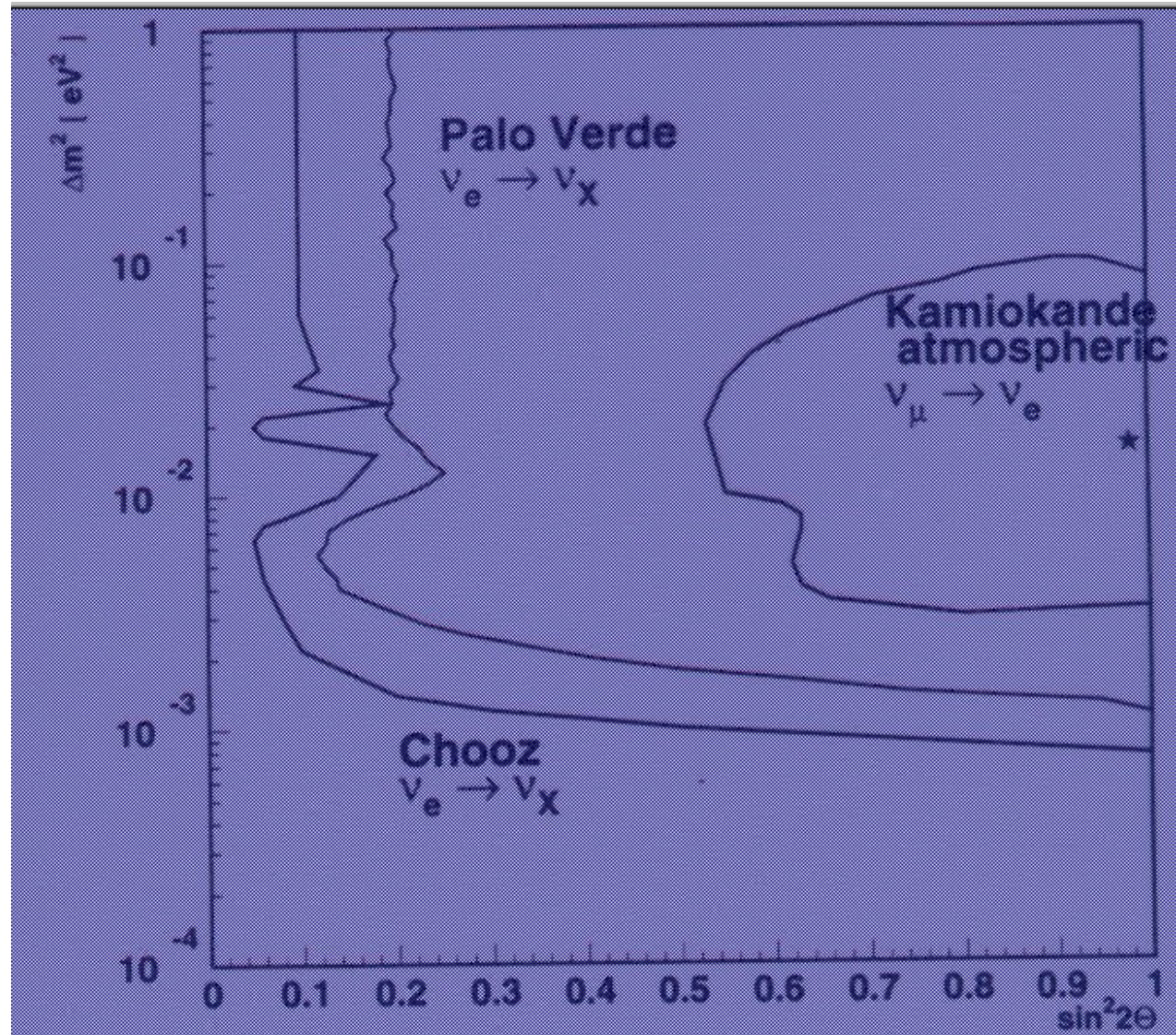
Events (day ⁻¹)	1998 ON	1998 OFF (890 m)	1999 ON	1999 OFF (750 m)
N_{cand}	38.2 ± 1.0	32.2 ± 1.0	52.9 ± 0.7	43.9 ± 1.4
b	19.5 ± 1.7		26.3 ± 2.2	
N_{detected}	18.7 ± 2.0	12.7 ± 2.0	26.6 ± 2.3	17.6 ± 2.6
R_{obs}	225 ± 24	140 ± 22	216 ± 19	140 ± 21
R_{calc}	218	155	218	130
<i>efficiency</i>	0.075	0.077	0.112	0.111

$R = 1.04 + 0.03 \text{ (stat)} + 0.08 \text{ (sys)} \Rightarrow$ No oscillation

Run till
end Summer 2000

2 new reduced power
periods

Not likely to do better
than Chooz



Three neutrinos families analysis

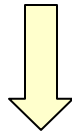
Reactor experiments

exclude two-family $\nu_\mu \rightarrow \nu_e$ oscillation

in parameters region where

ν_μ deficit in atmospheric experiments

favours two-family $\nu_\mu \rightarrow \nu_\tau$ (or ν_s)



(at least) 3-flavour analysis

3-flavour mixing parametrization

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} \quad \left. \begin{aligned} \nu_\alpha &= \sum_{k=1}^3 U_{\alpha k} \nu_k \\ \sum_{k=1}^3 |U_{\alpha k}|^2 &= 1 \\ \sum_{k,k' \neq k}^{=1,3} \Delta m_{kk'}^2 &= 0 \end{aligned} \right\} \alpha = e, \mu, \tau$$

6 (8) parameters $\left\{ \begin{array}{l} 2 \Delta m_{kk'}^2, \\ 3 |U_{\alpha k}| \end{array} \right.$
 Dirac (Majorana) $\left\{ \begin{array}{l} 1 \text{ (3) phases} \end{array} \right.$

CKM-like matrix standard parametrization)

$$U = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ & s_{23}c_{13} & \\ & c_{23}c_{13} & \end{pmatrix} \quad \left\{ \begin{array}{l} c_{ij} = \cos \theta_{ij} \\ s_{ij} = \sin \theta_{ij} \end{array} \right\} \quad i, j = 1, 2, 3 \text{ generation numbers}$$

3-family flavour at the strong mass hierarchy approximation

if $m_3 \gg m_1, m_2$

$\Delta m^2 = m_3^2 - m_1^2 \approx m_3^2 - m_2^2$ e.g. 10^{-3} eV^2 atmospheric neutrinos

$\delta m^2 = m_2^2 - m_1^2$ e.g. 10^{-6} eV^2 solar neutrinos

$\Delta m^2 \gg \delta m^2$



L/E region where $\Delta m^2 E / L$ causes oscillation

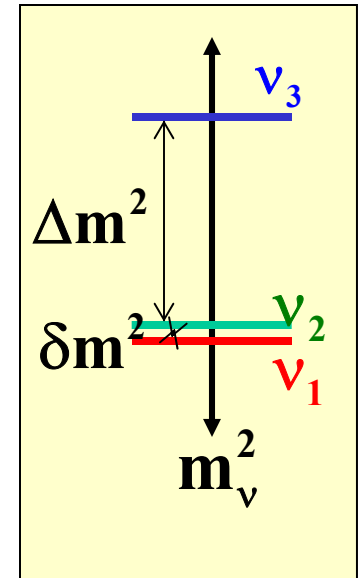
(e.g. atmospheric neutrinos $\Delta m^2 = 10^{-3} \text{ eV}^2$, $E = 1 \text{ GeV}$, $L = 1000 \text{ km}$)

and $\delta m^2 E / L \approx 0$



$$P(\nu_\alpha \rightarrow \nu_{\beta \neq \alpha}) \approx 4 |U_{\alpha 3} U_{\beta 3}|^2 \sin^2(1.27 \Delta m^2 E / L)$$

$$\sin^2 2\theta_{\alpha\beta}^{\text{eff}} = 4 |U_{\alpha 3} U_{\beta 3}|^2$$



Physics governed by:

- Δm^2
- flavour contents of ν_3
- effective 2-flavour like oscillation

Effective 2-family atmospheric ν_μ disappearance in 3-family mixing

$$\sin^2 2\theta_{\mu e}^{\text{eff}} = 4 |U_{\mu 3}^2 U_{e 3}^2| = \sin^2 2\theta_{13} \sin^2 \theta_{23}$$

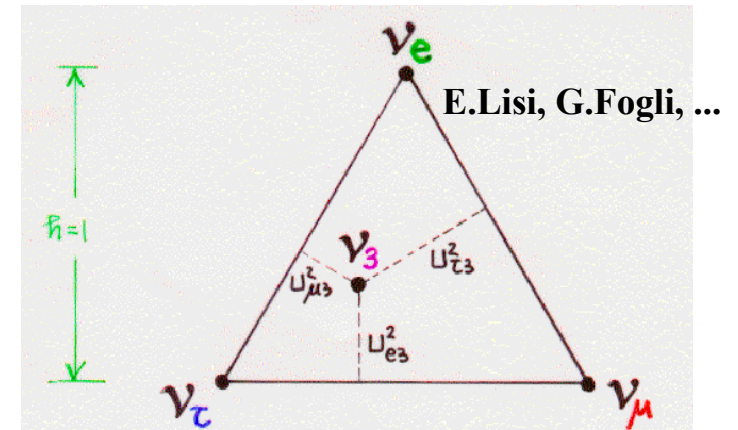
$$\sin^2 2\theta_{\mu \tau}^{\text{eff}} = 4 |U_{\mu 3}^2 U_{\tau 3}^2| = \sin^2 2\theta_{23} \cos^4 \theta_{13}$$

$U_{e 3}^2$ small (reactors)

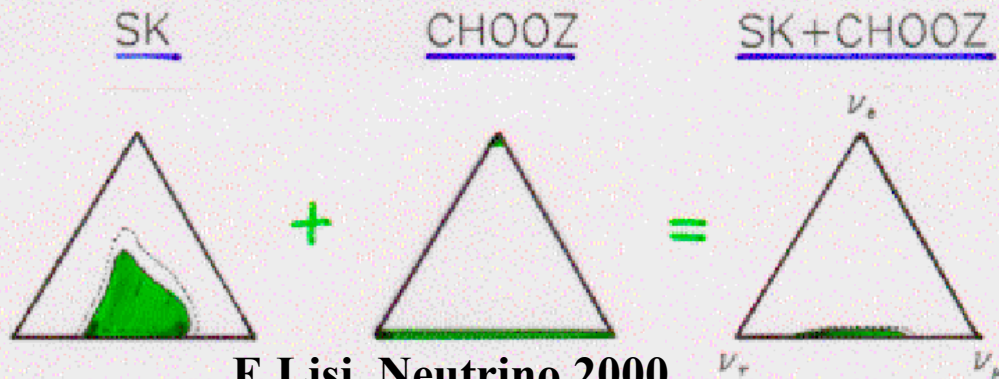
$4 |U_{\mu 3}^2 U_{\tau 3}^2| \approx 1$ (full mixing atmospheric)

$$U_{\mu 3}^2 + U_{\tau 3}^2 \approx 1$$

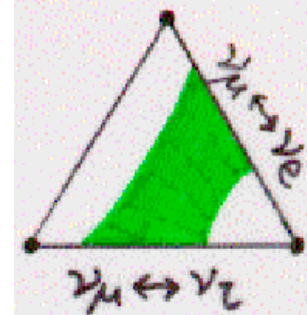
$$\left. \begin{array}{l} U_{e 3}^2 \text{ small (reactors)} \\ 4 |U_{\mu 3}^2 U_{\tau 3}^2| \approx 1 \text{ (full mixing atmospheric)} \\ U_{\mu 3}^2 + U_{\tau 3}^2 \approx 1 \end{array} \right\} \Rightarrow \left\{ \begin{array}{l} U_{e 3}^2 < 0.1 ? \\ U_{\mu 3}^2 \approx U_{\tau 3}^2 \approx 0.5 ? \end{array} \right.$$



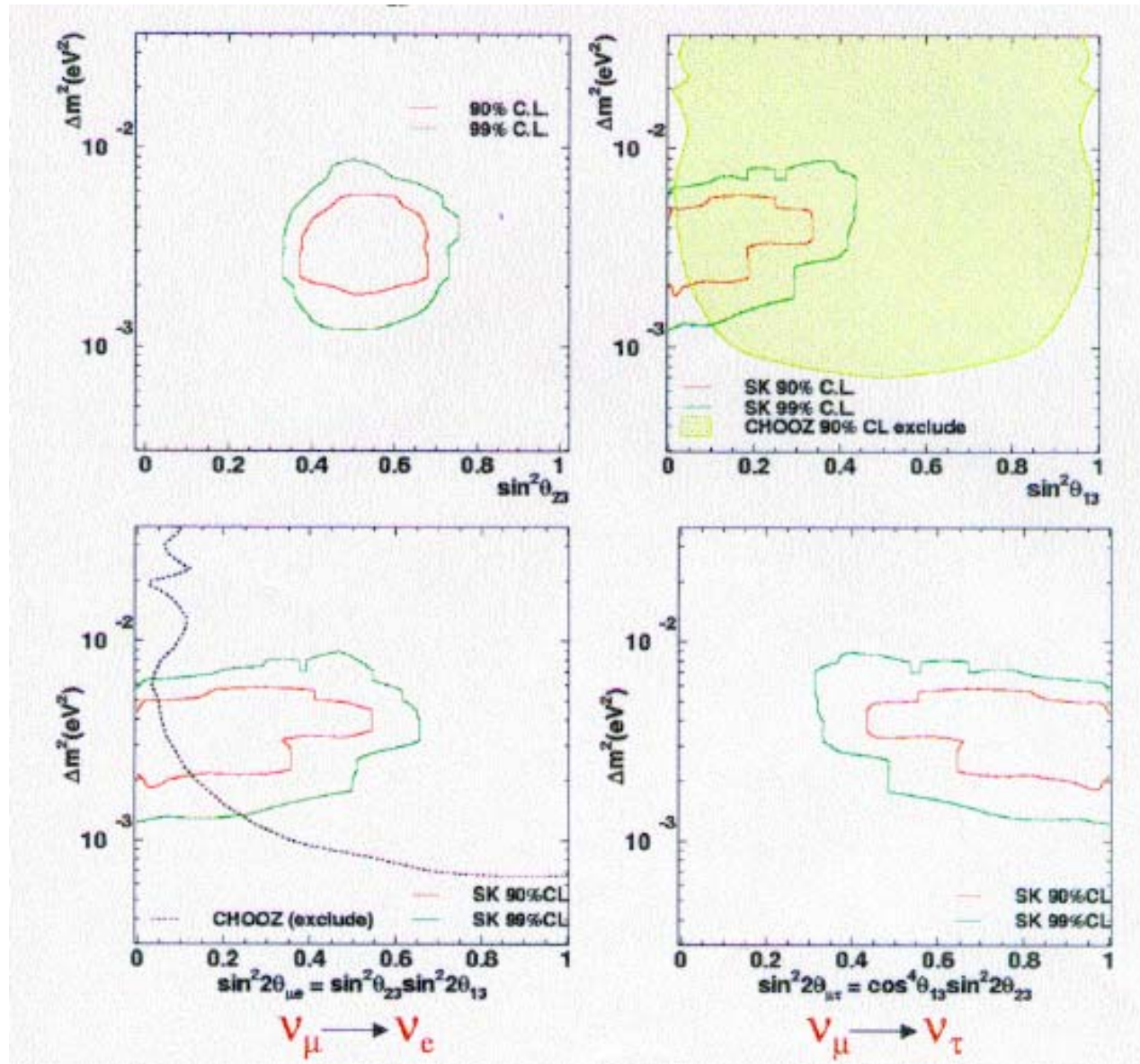
● SK + CHOOZ '99 data analysis :



● Compare with situation < 1998 :



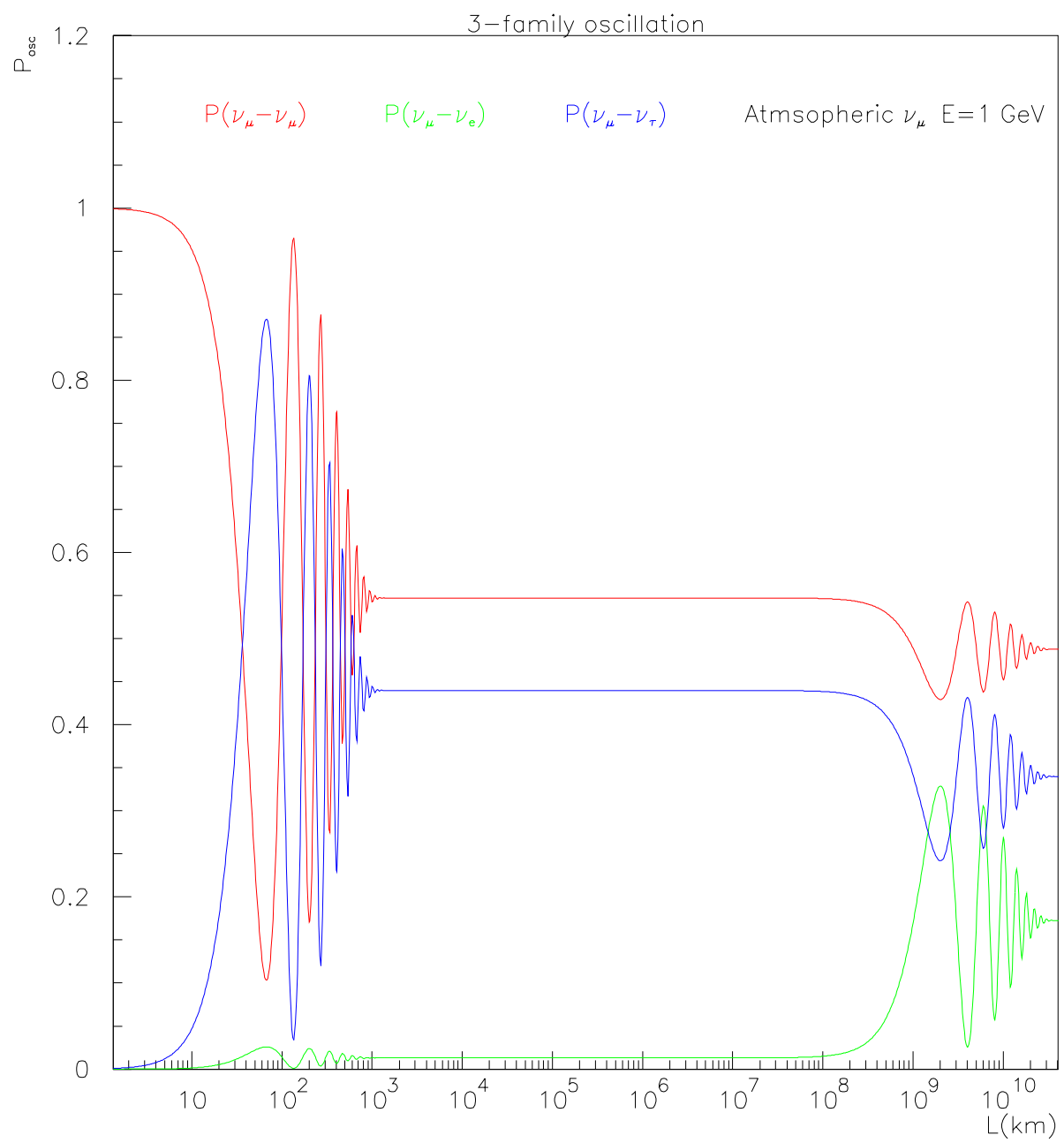
H.Sobel Neutrino2000



Space is left
for $U^2_{e3} \neq 0$

Conclusions:

- **No evidence for ν_e disappearance in LBL reactor experiments**
 - **Reactor + Atmospheric neutrino experiments**
 - + **in 3-flavour strong mass hierarchy model**
- room left for a small ν_e contents in ν_3**
- **No more constraining data to be expected from reactors in near future**



Neutrino Oscillation Experiments at Low Energy Accelerators (Beam Stoppers)

Search for $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillation at rather large $\Delta m^2 > \sim 0.1 \text{eV}^2$

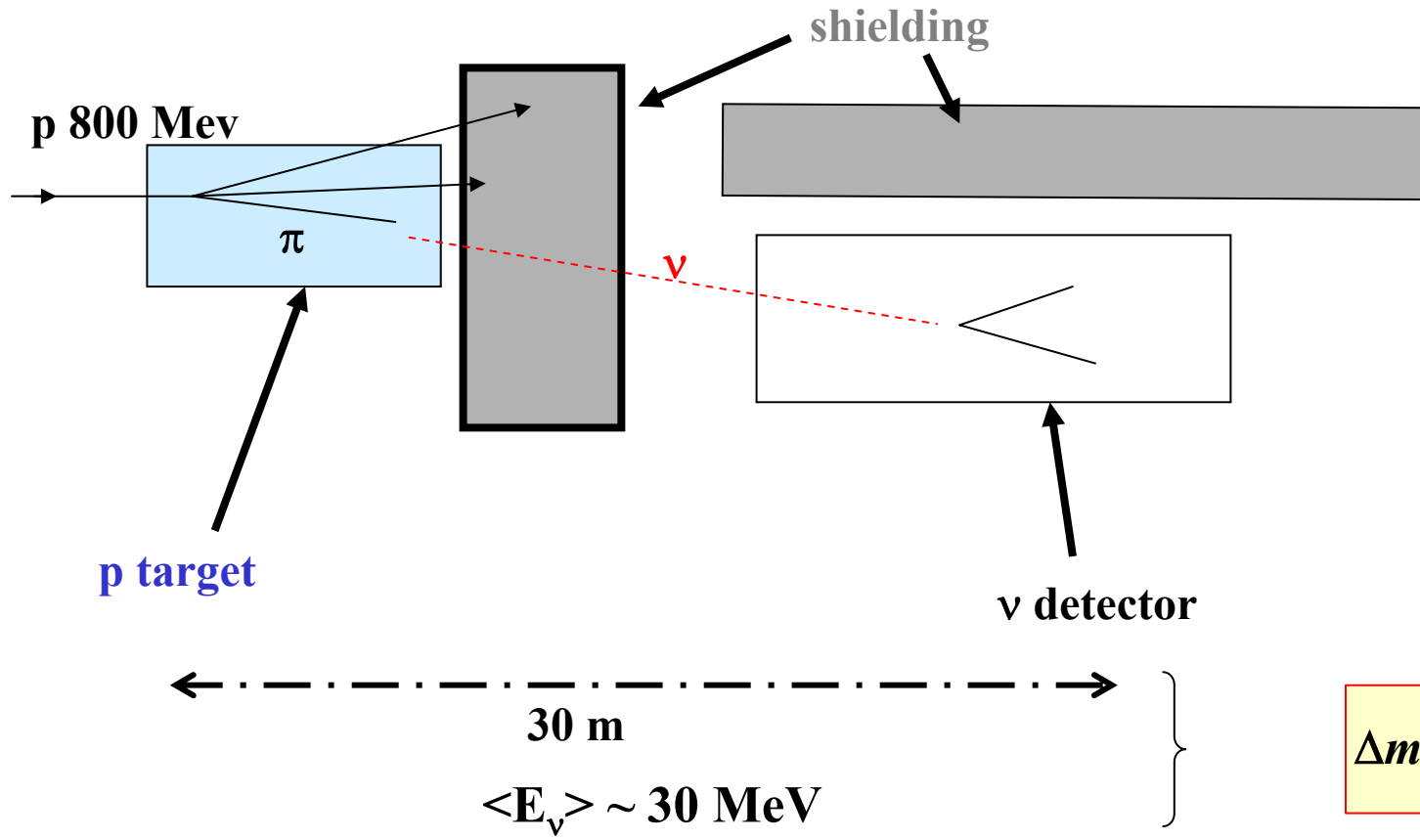
Compare somehow conflicting results from two similar experiments:

KARMEN2: no signal

LSND: statistically significant signal

LSND: G.Mills Neutrino 2000
KARMEN2: K.Eitel

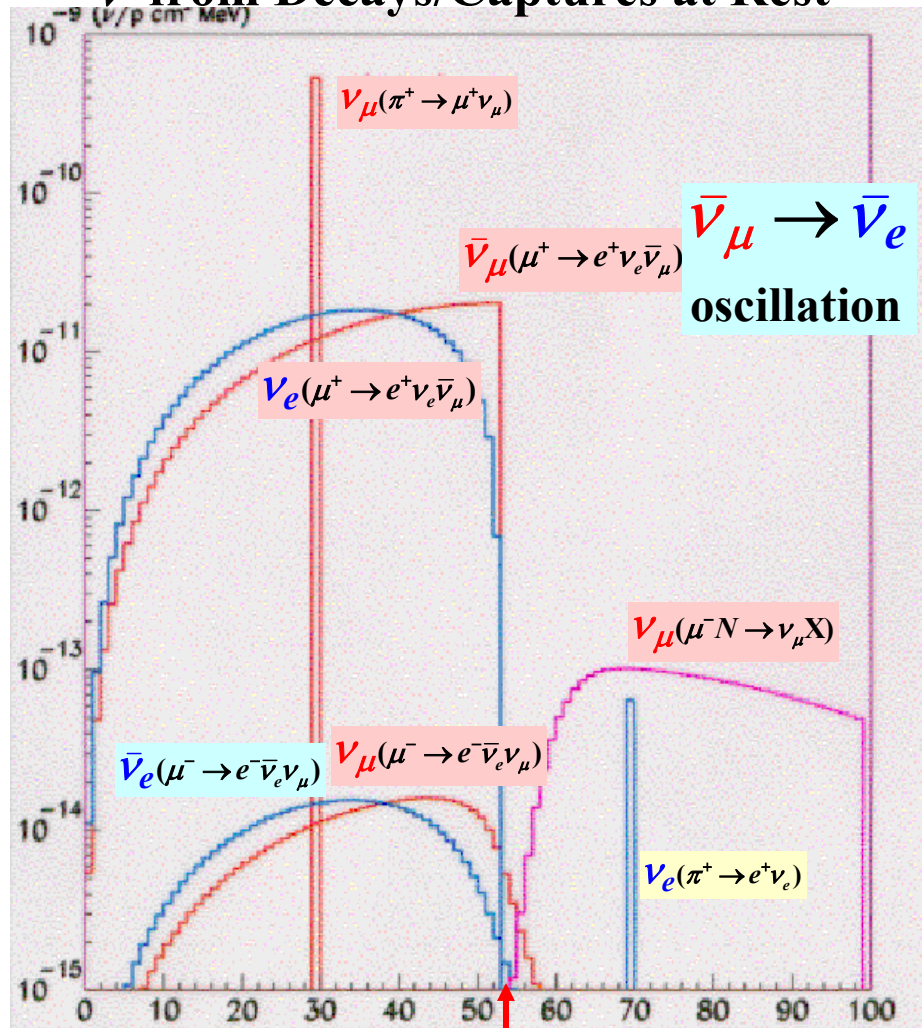
Conceptual design:



$$\Delta m^2 \approx \frac{L}{E} \approx 1 \text{ eV}^2$$

The ν Energy spectra

ν from Decays/Captures at Rest

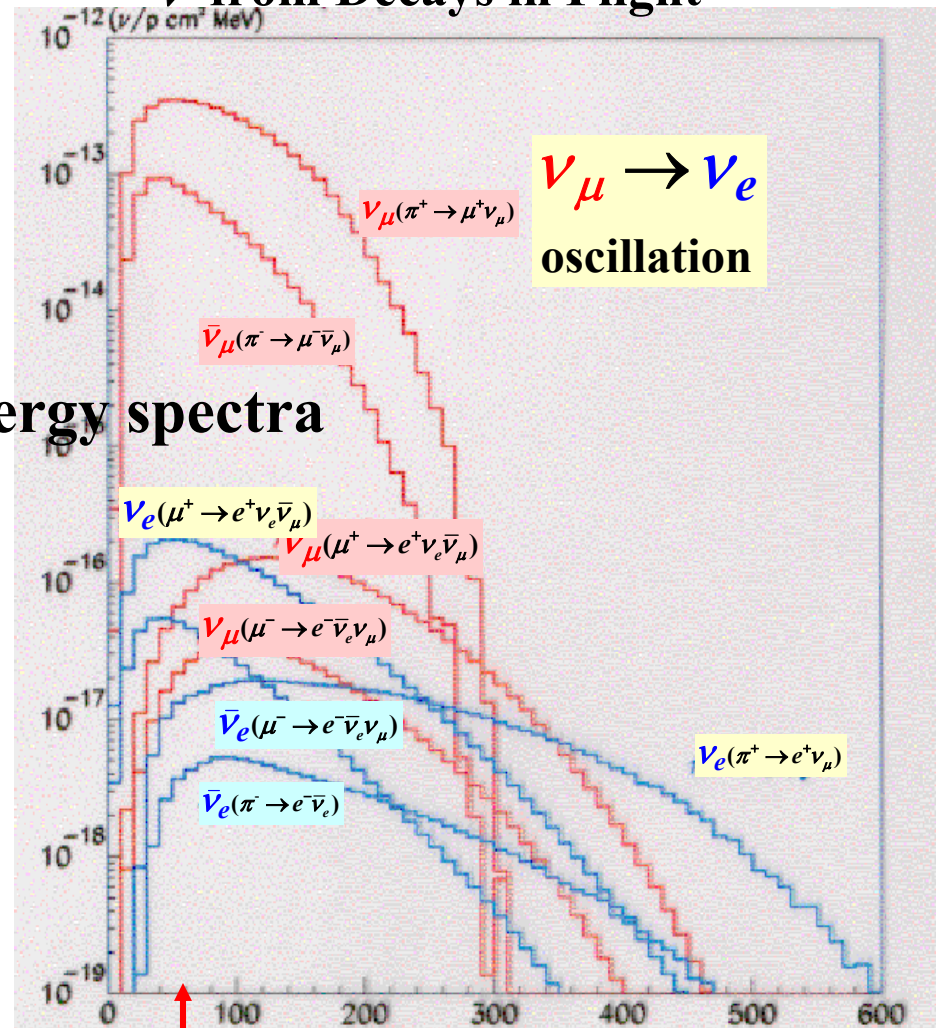


$\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
 $E_\nu < 53 \text{ MeV}$

53 MeV

spectra are for LSND

ν from Decays in Flight



$\nu_\mu \rightarrow \nu_e$
 oscillation

53 MeV

$\nu_\mu \rightarrow \nu_e$
 $E_\nu > 53 \text{ MeV}$

Detectors

vessel filled with oil + liquid scintillator
doped with neutronphage

several light signal by arrays of PMT's

Signal

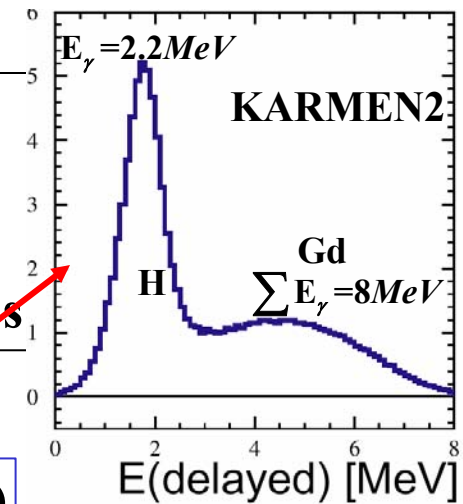
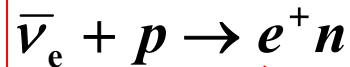
$$E_{\bar{\nu}_e}^{thresh} = 1.8 \text{ MeV}$$
$$< E_{e^+} > \approx 3 \text{ MeV}$$

scintillation γ 's

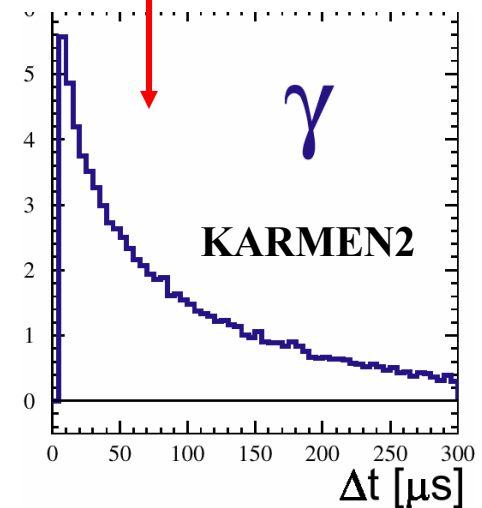
$e^+ - e^-$ annihilation : 2 γ of 0.511 keV

Cerenkov light

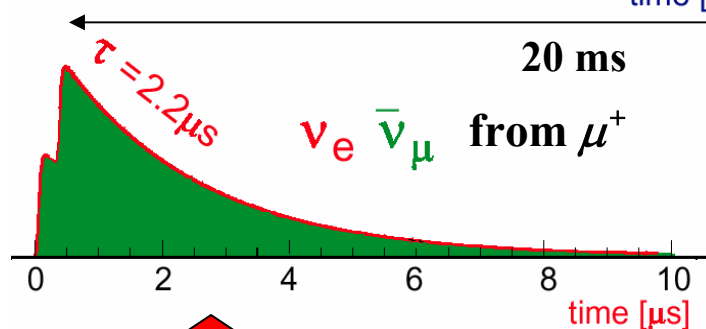
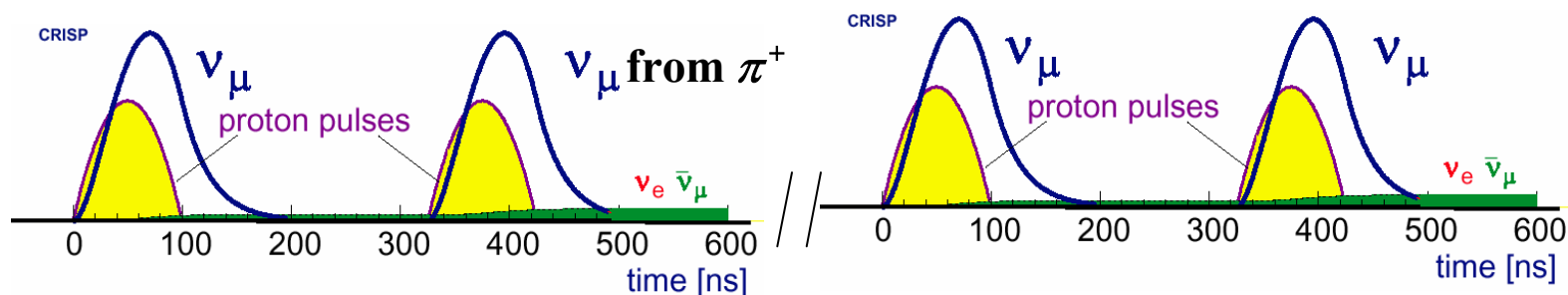
nuclear capture $\rightarrow \gamma$'s ($\sum E_\gamma$ known)



Space and
delayed time
correlation



Main Karmen2 pro: Time structure of ISIS p source



and
prompt e^+ signal from QE $\bar{\nu}_e p$
from $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillation

- ν_μ background measured in $[t_0, t_0 + 600 ns]$
- $\bar{\nu}_\mu$ and $\bar{\nu}_e$ from $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ signal recorded in $[t_0 + 600 ns, t_0 + 10.6 \mu s]$ with $2.2 \mu s$ slope
- cosmic background measured in $[t_0 + 10.6 \mu s, t_0 + 20 ms]$
- small duty cycle \rightarrow small cosmic background

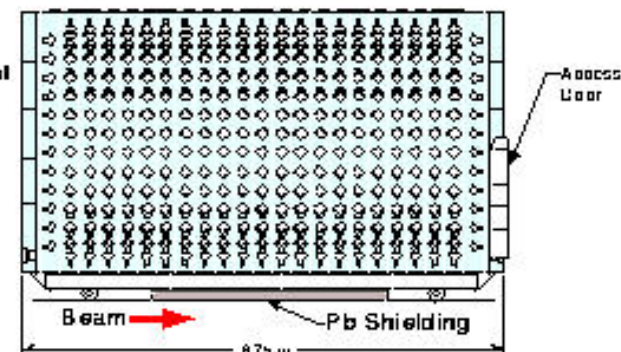
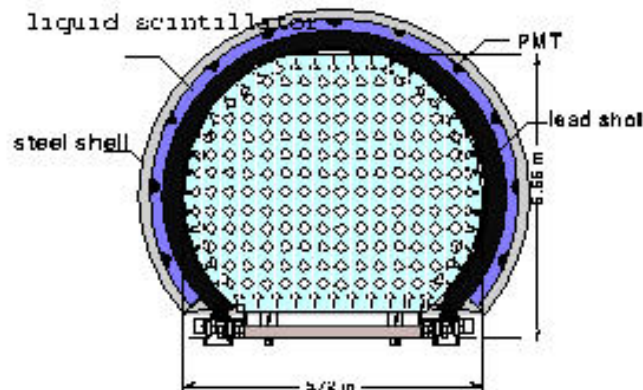
- Segmented detector: better $n-e$ space correlation
- High scintillator concentration: $\times 4$ better E resolution

Main LSND pro: Electron ID and direction

**Homogeneous detector + low scintillator concentration
⇒ Cerenkov ring as e^+ signature**

- **×3 Mass**
- **$L=18\text{m}$ (instead of 30m) ⇒ lower Δm^2**
- **3% of DIF $\pi^+ \rightarrow \mu^+ \nu_\mu$ (instead of 0.1%)**
⇒ **higher energy beam component**
⇒ **$\nu_\mu \rightarrow \nu_e$ oscillation via $\nu_e C \rightarrow e^- N$**

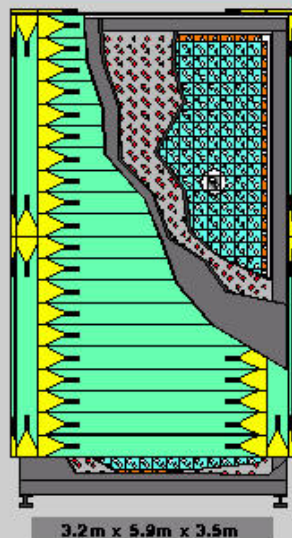
LSND detector at LAMPF, Los Alamos



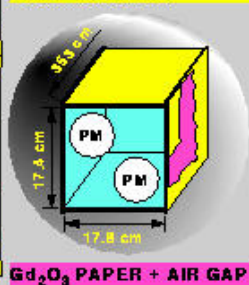
KARMEN DETECTOR

56t liquid scintillator calorimeter

96% active volume



OUTER VETO 1 NE110
PASSIVE SHIELD 20 cm
INNER VETO 98
SEGMENTED 18x32
CENTRAL DETECTOR
3" PHOTOMULTIPLIER
ACRYLIC GLASS



$$\Delta E/E = \frac{11.5\%}{\sqrt{E(\text{MeV})}} \quad (\Leftrightarrow 80 \text{ pe/MeV})$$

$$\sigma t = 0.4 \text{ ns} \quad \sigma x = 6.0 \text{ cm}$$

KARMEN-II detector at ISIS, RAL

Statistical analysis difficulties

KARMEN-2

- **no signal and very low expected background : place an Upper Limit**
- **(for long: 0-event observed sample, 3 expected background)**
- **non physical max likelihood : $\sin^2 2\theta < 0$**

LSND

- **signal region in parameter space computed from rapidly oscillating likelihood function with many local maxima**

Profusion of recent papers and workshops

on our to fix C.L. limits from likelihood functions

(starting G.J.Feldman & R.D.Cousins, Phys Rev D57(1998)3873)

KARMEN-2 results

data set after final cuts

11 candidates

3.94 \pm 0.51 ■ ν_e -induced CC sequ.

3.52 \pm 0.30 ■ ν -induced random bg.

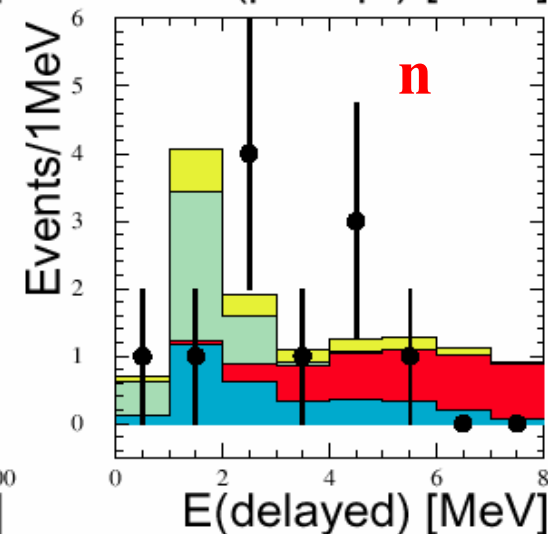
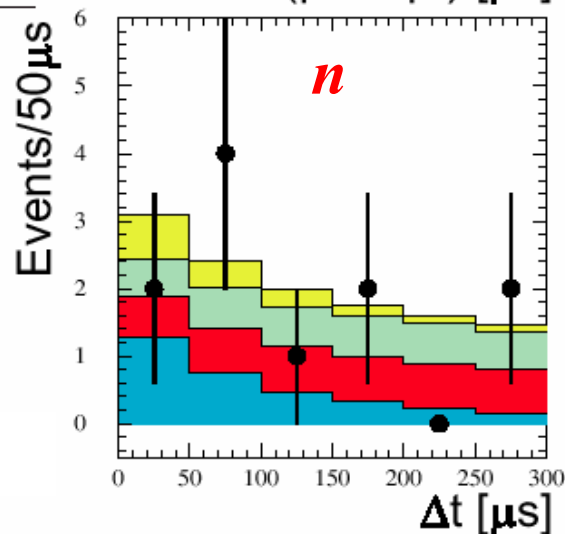
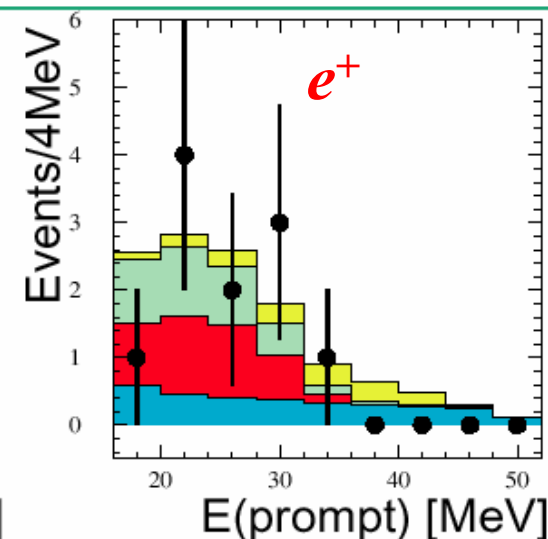
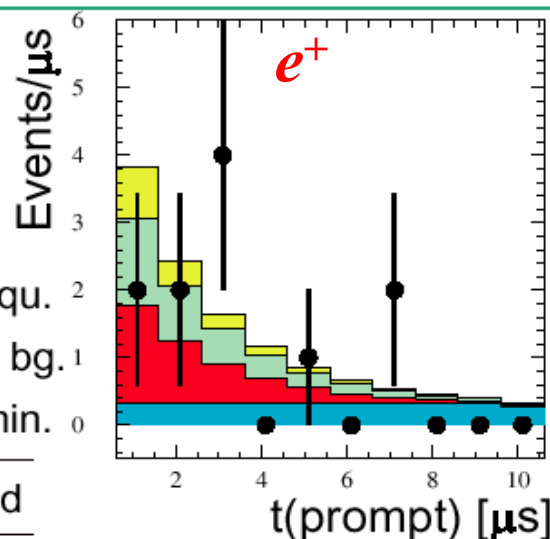
1.67 \pm 0.17 ■ $\bar{\nu}_e$ intrinsic contamin.

3.17 \pm 0.17 ■ cosmic background

12.29 \pm 0.63 total background

no osci signal

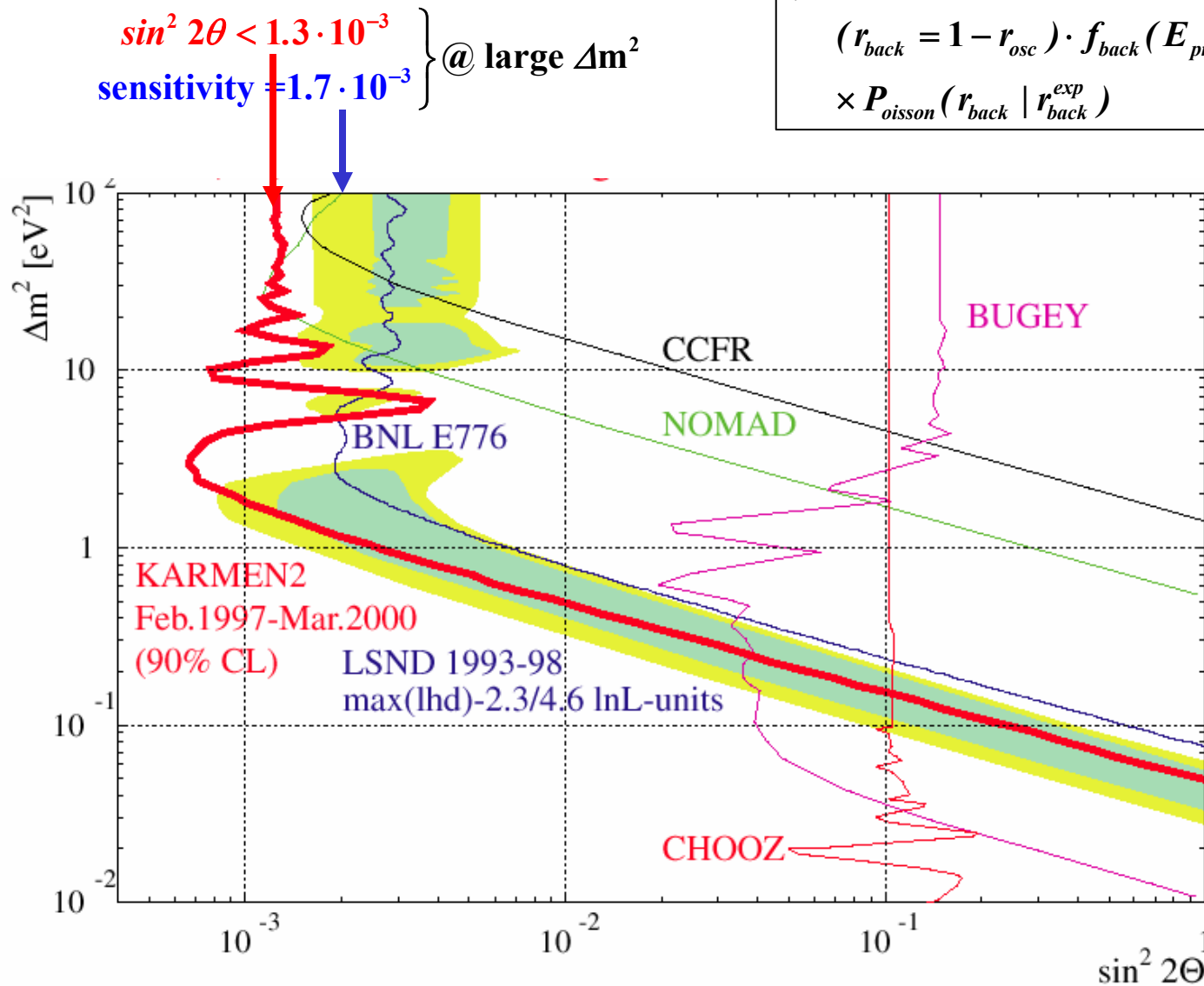
All backgrounds
measured except
intrinsic



KARMEN-2 exclusion plot

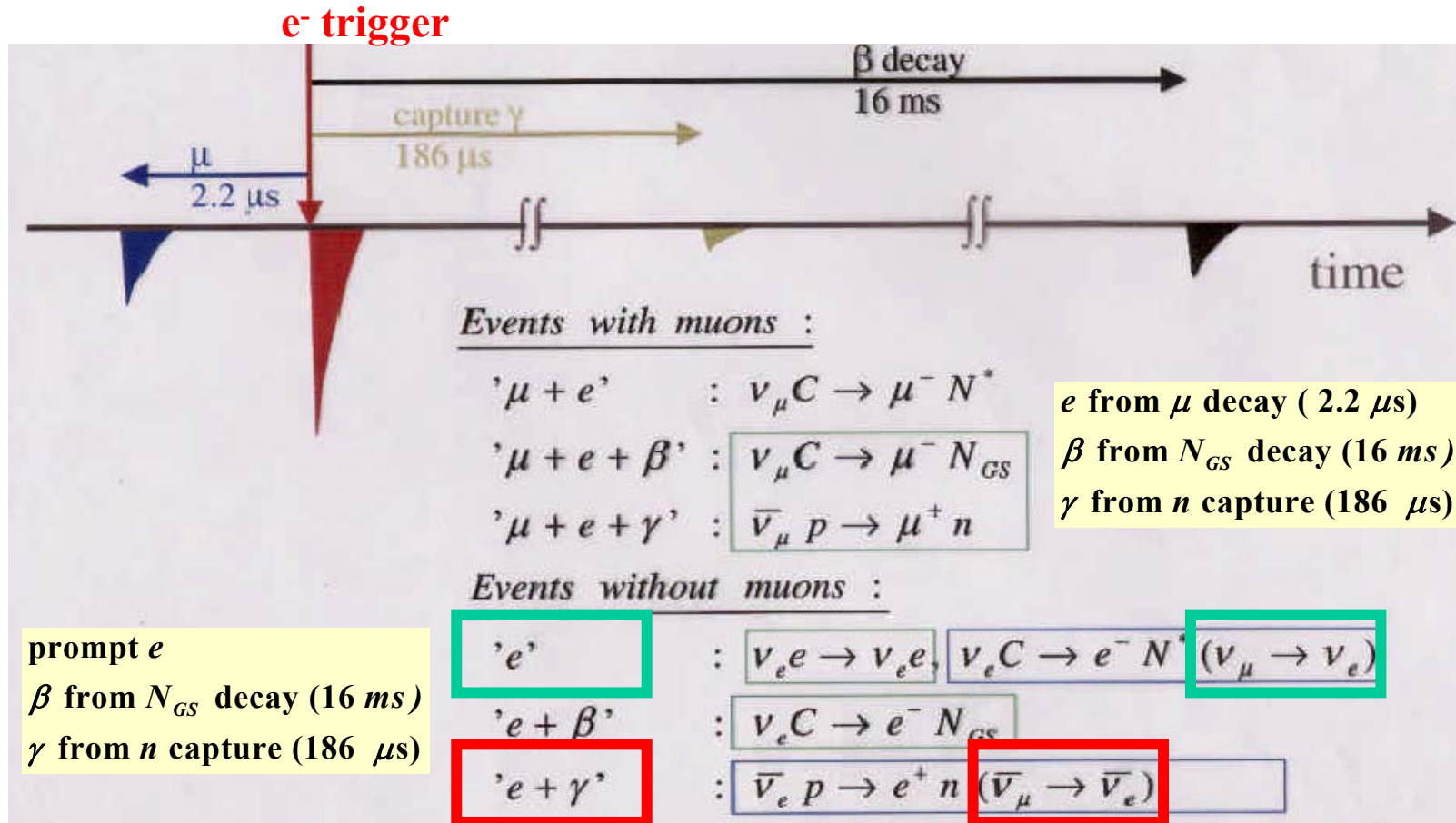
$$L(\sin^2 2\theta, \Delta m^2) =$$

$$\prod_{k=1}^{11} \{ r_{osc} \cdot f_{osc}(E_{prompt}, E_{delayed}, t_{prompt}, \Delta t, \Delta \vec{r})_k + \\ (r_{back} = 1 - r_{osc}) \cdot f_{back}(E_{prompt}, E_{delayed}, t_{prompt}, \Delta t, \Delta \vec{r})_k \} \\ \times P_{oisson}(r_{back} | r_{back}^{exp})$$



+ Unified frequentist approach (F.-C.)

New LSND global analysis of all event categories with a common Electron trigger and E_e in [20-200] MeV

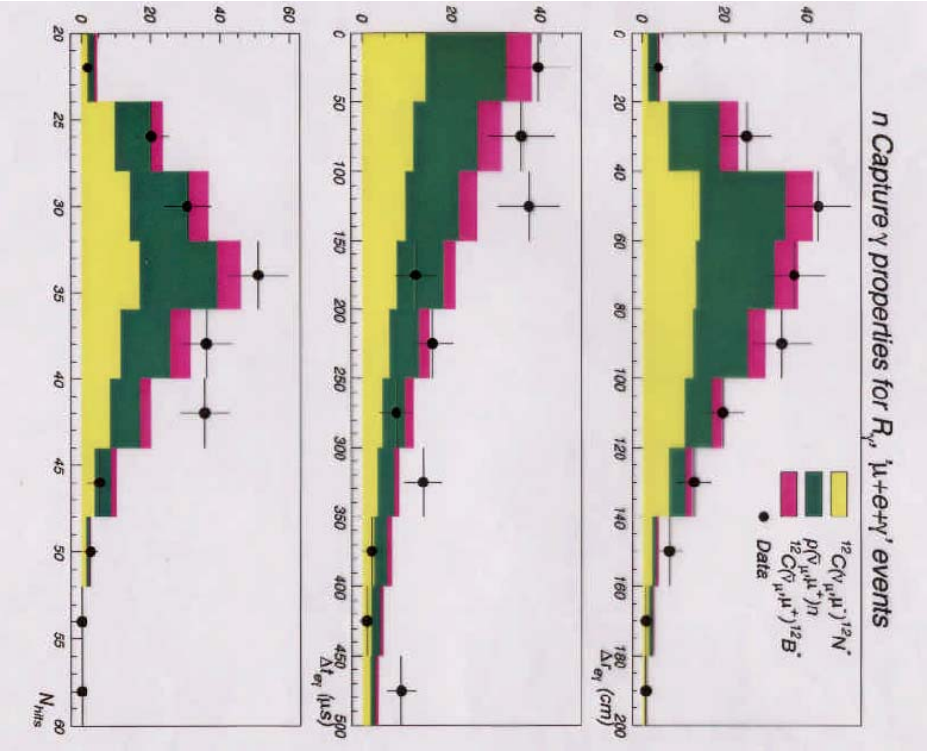


New LSND global analysis of all event categories with a common Electron trigger and E_e in [20-200] MeV

Global fit

- ◆ to all relevant distributions
 - $E(e, \beta, \gamma, \mu)$
 - $\Delta t(e - \beta, \gamma, \mu)$
 - $\Delta r(e - \beta, \gamma, \mu)$
 - $\theta(\nu - e^-)$
 - R : ratio of likelihood of prompt (e^-) and delayed events (γ) to be correlated/accidental
- ◆ for all electron trigger events categories
- ◆ with parameters:
 - π^+/π^- production ratio
 - all DAR and DIF π and μ
 - efficiencies μ, e, β, γ

Oscillation signal : “ $e \gamma$ ” events with large R

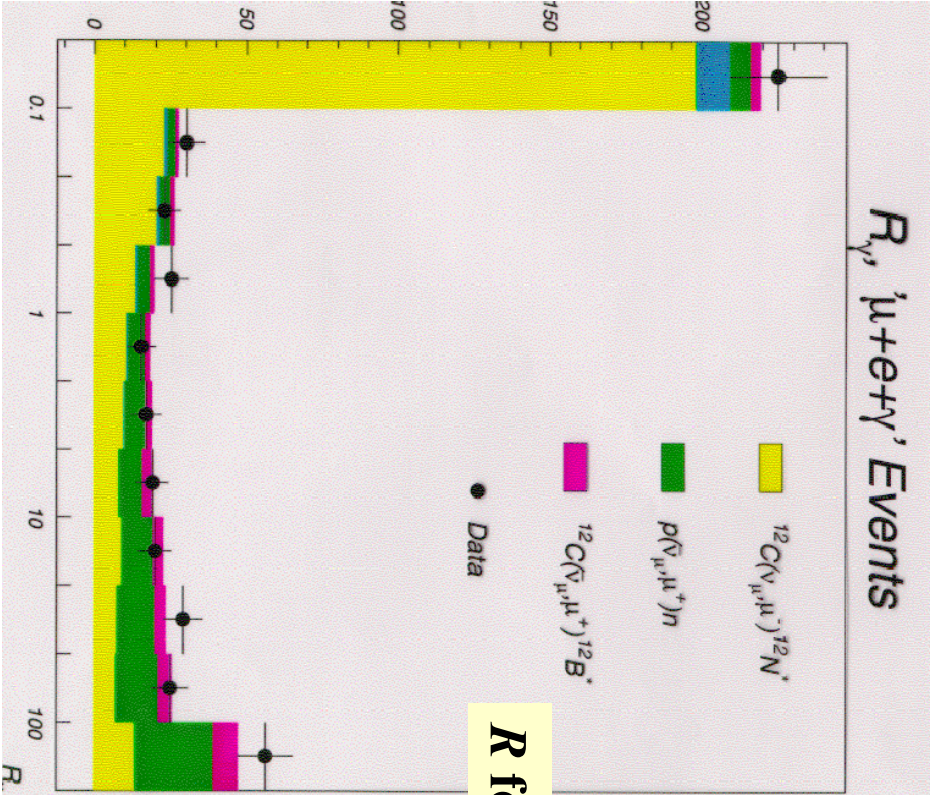


Variables entering in R

$$\Delta r \text{ (} e - \gamma \text{)}$$

$$\Delta t \text{ (} e - \gamma \text{)}$$

$$\text{PMT hits} \div E$$



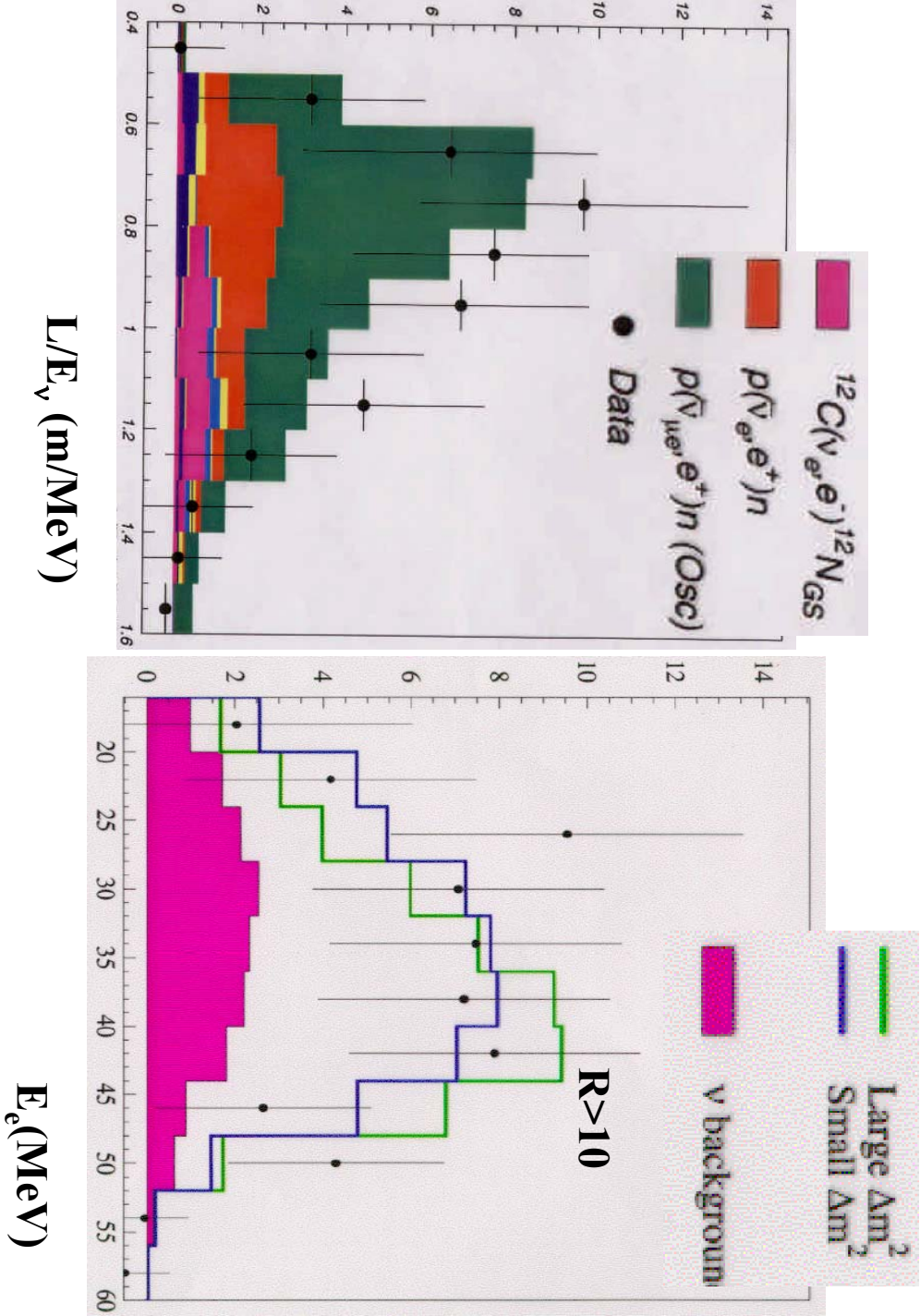
R for no-oscillation channel

The oscillation signal in E_e in [20-60] MeV

Excess of events at $R > 10$ (large e- γ correlation likelihood)

beam on total	beam off background	expected ν background	excess of events
83	33.7	16.6	32.7 \pm 9.2

Event excess is E_e dependent



Fit to full R distribution

$$P_{\text{osc}} = 0.0025 \pm 0.0006 \pm 0.0004$$

LSND signal region

Relaxing cuts:

- E_e in [20-200] MeV
- $R > 0$

+

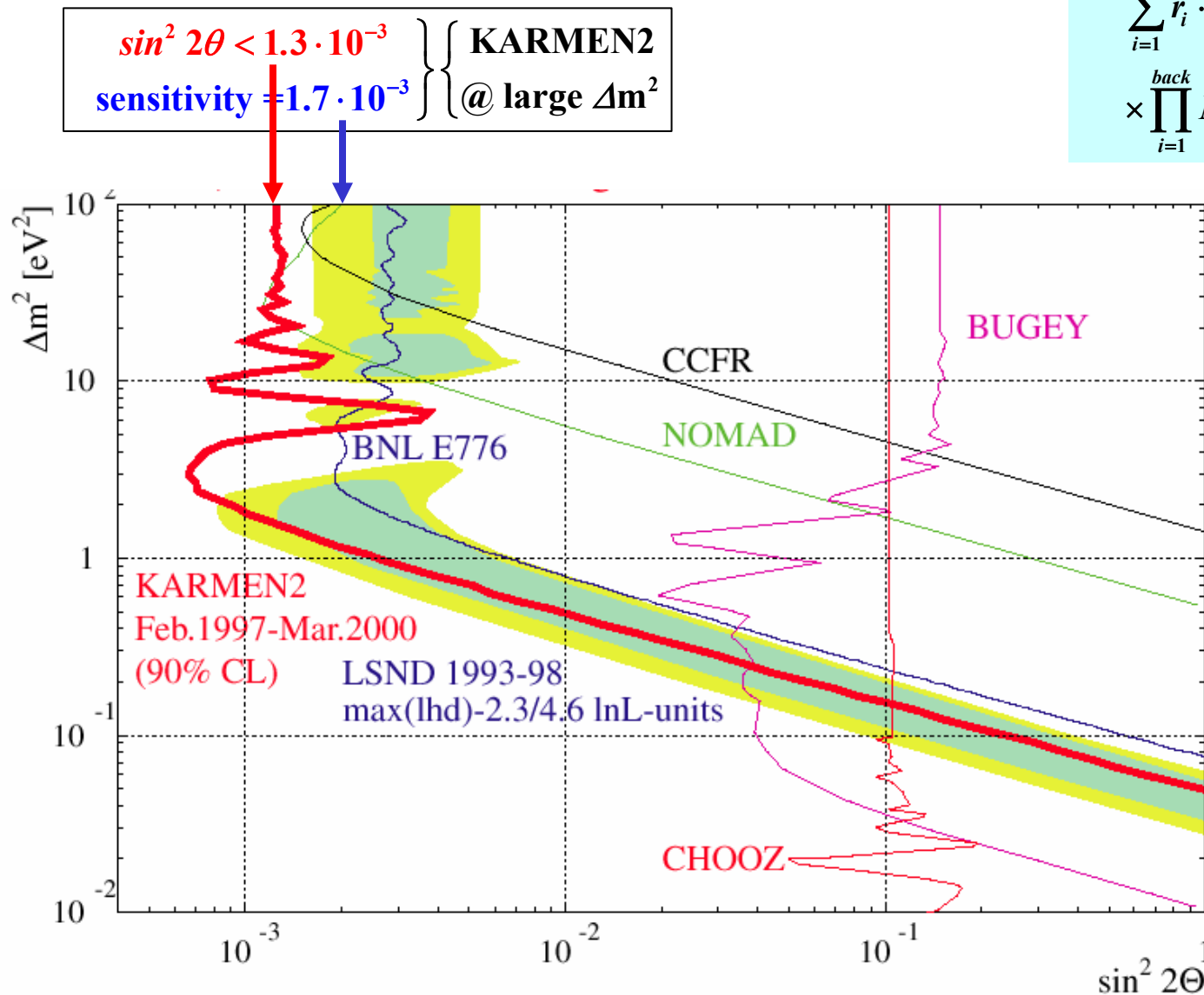
$$L(\sin^2 2\theta, \Delta m^2) =$$

$$\prod_{k=1}^{N_{event}} \{ r_{osc} \cdot f_{osc}(E_e, R, L_\nu, \cos \theta_\nu)_k + \sum_{i=1}^{back} r_i \cdot f_i(E_e, R, L_\nu, \cos \theta_\nu)_k \} \times \prod_{i=1}^{back} N_{normal}(r_i \cdot N | N_i^{exp})$$

+

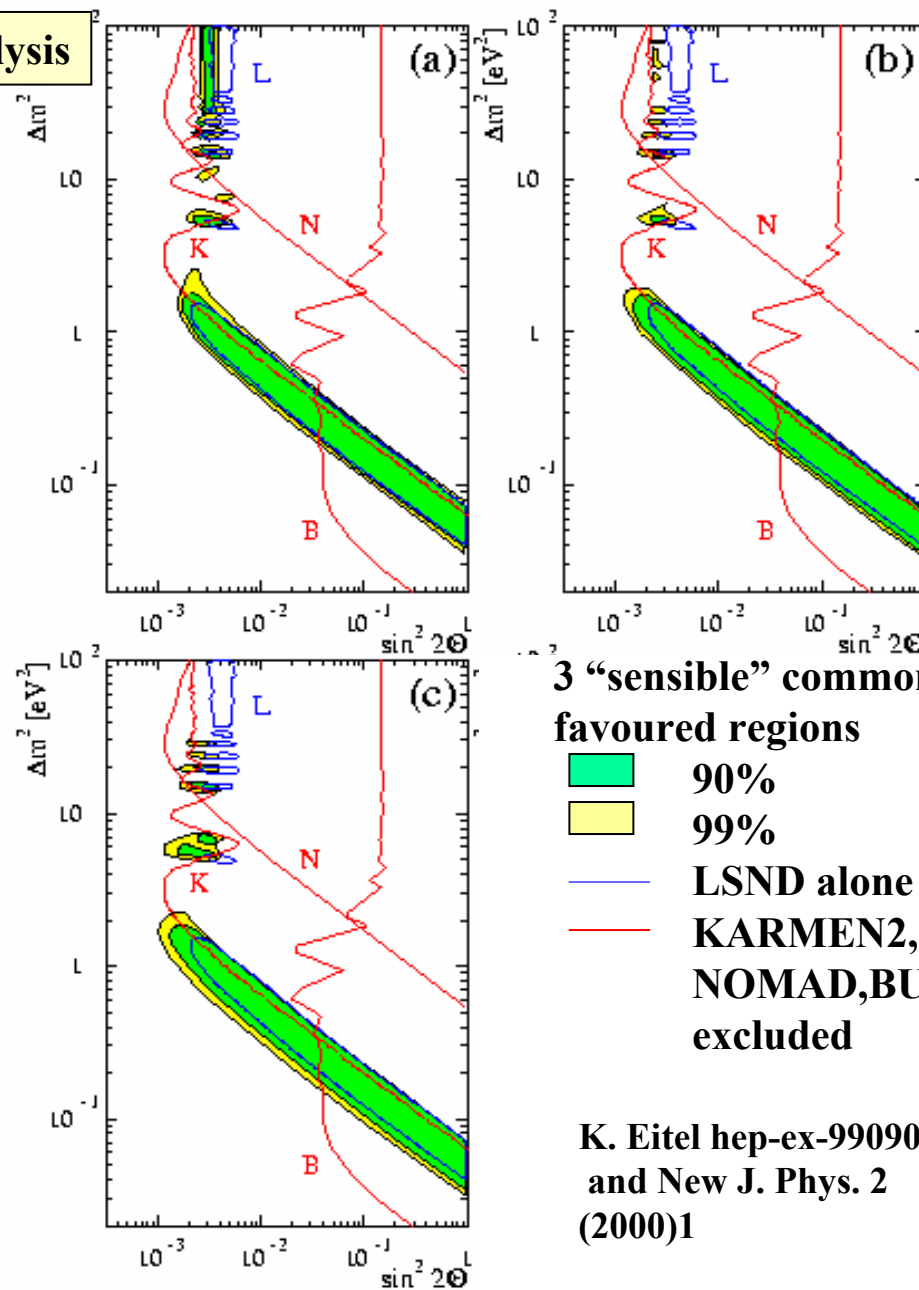
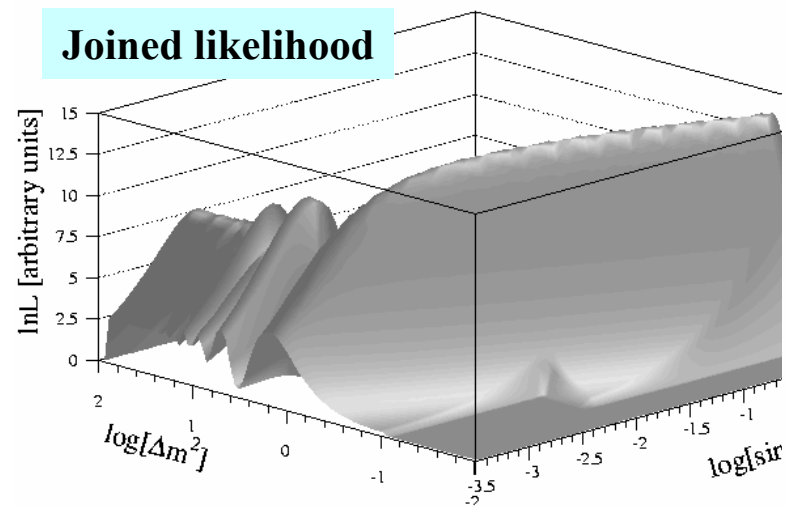
Cut at $\Delta L =$
2.3 (90%)
4.6 (99%)

Compatible
1993-95 &
1996-98
signals



Preliminary joined KARMEN2-LSND analysis

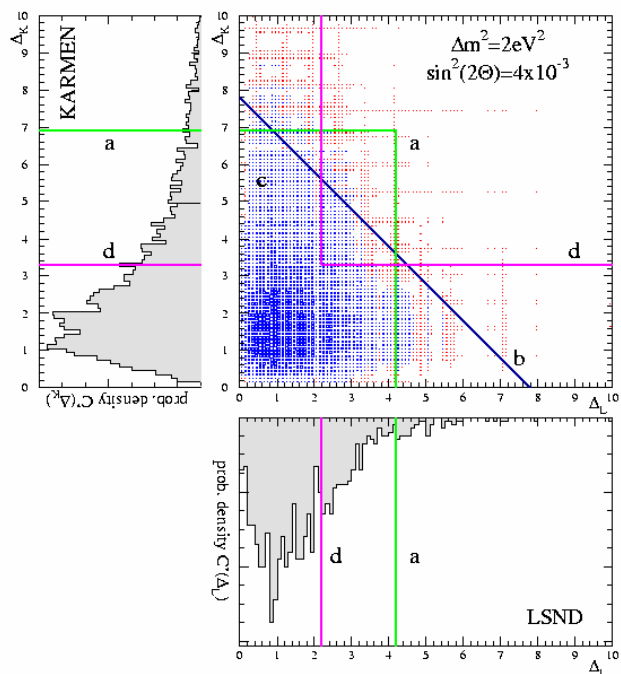
Joined likelihood



3 “sensible” common
favoured regions

- 90%
- 99%
- LSND alone
- KARMEN2,
NOMAD,BUGEY
excluded

K. Eitel hep-ex-990906
and New J. Phys. 2
(2000)1



Conclusions

- LSND signal in $\Delta m^2 \sim \text{eV}^2$ is one of the 3 oscillation signals
- No evidence that result is wrong
- Allowed LSND parameters space domain will not be fully covered by KARMEN2 (\Rightarrow spring 2001): not enough statistics given background
- Need for a joined analysis based on a common likelihood function based on the final data,
- Need new experiment(s) with higher sensitivity:
MiniBOONE approved at FERMILAB from 2001

I216 proposal at CERN PS 2001 

In the mean time
either

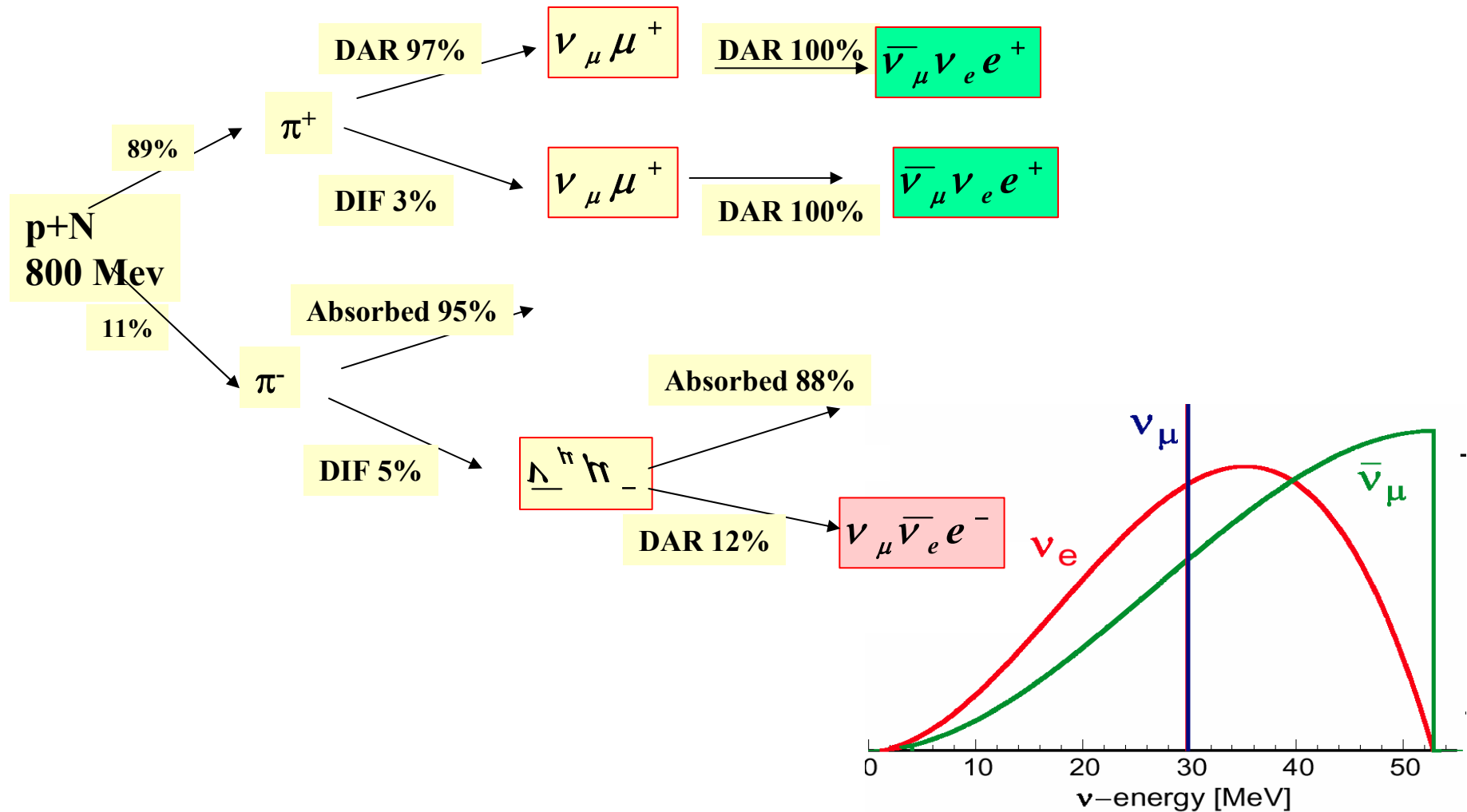
$$3 \Delta m^2$$

$$\nu_e \nu_\mu \nu_\tau \nu_s$$

or



The neutrino production chain (numbers are for LSND)



Search for $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$ oscillations
from $\mu^+ \rightarrow e^+$ decays at rest

Neutrino Oscillation Experiments at High Energy Accelerators

CHORUS and NOMAD short baseline experiment

Search for $\nu_\mu - \nu_\tau$ oscillation

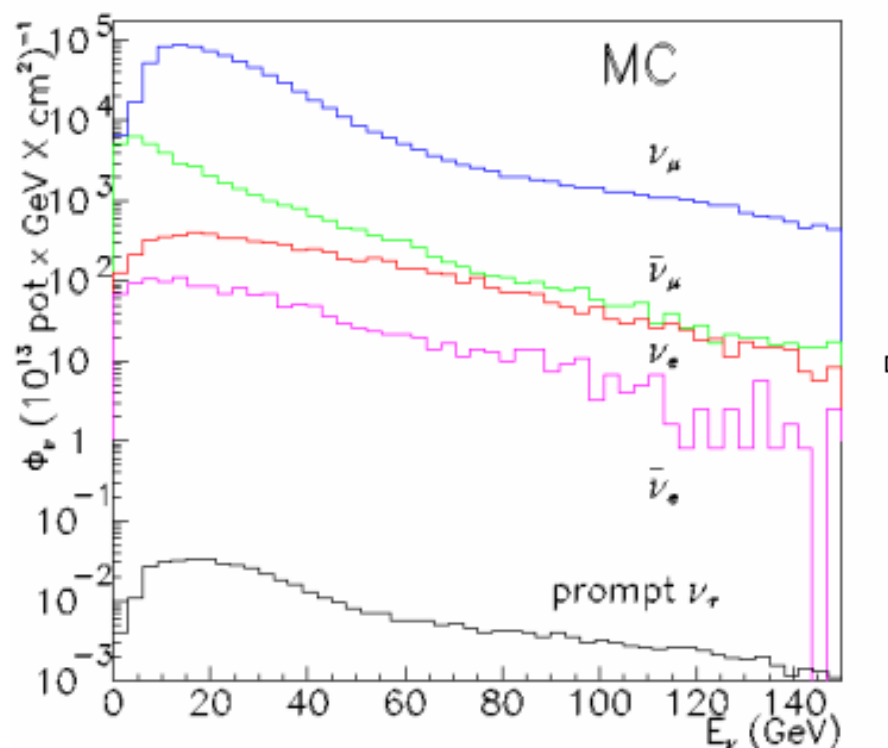
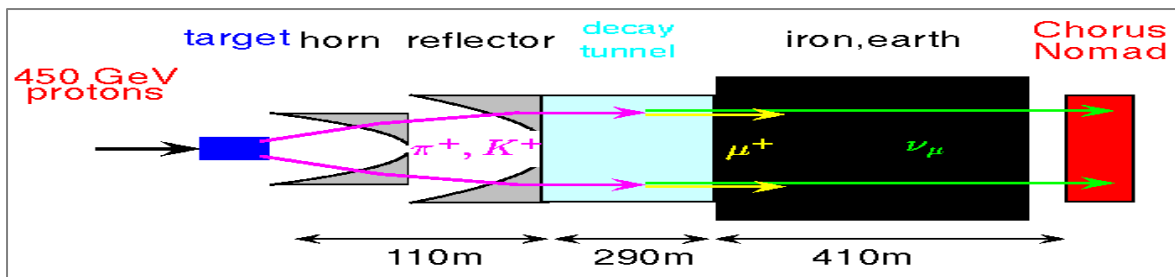
ν_τ appearance
in same high energy ν_τ free ν_μ beam
at the CERN SPS Wide Band Neutrino Beam

Motivation (early 1990's)

Search for “hot dark matter” candidates with $m_\nu > 1 \text{ eV}$

with ~ 50 times better sensitivity than E531: $P_{osc}(\nu_\mu - \nu_\tau) > 10^{-4}$

CERN Wide Energy-Band Neutrino Beam Line



	Φ_{relative}	$\langle E_\nu \rangle$ [GeV]
ν_μ	1	27
$\bar{\nu}_\mu$	0.056	19
ν_e	0.009	~40
$\bar{\nu}_e$	0.002	~32
ν_τ	$3 \cdot 10^{-6}$	~43

Irreducible prompt ν_τ background from $D_S \rightarrow \tau \nu_\tau$
 less than 0.1 event in 4 years
 well below sensitivity

Maximum sensitivity @

$$\Delta m^2 \approx \frac{27 \text{ GeV}}{0.6 \text{ km}} \approx 50 \text{ eV}^2$$



AUSTRALIA Melbourne, Sydney

CERN

CROATIA Zagreb

FRANCE LAPP Annecy, Paris LPNHE, Saclay DAPNIA

GERMANY Dortmund.

ITALY Cosenza, Firenze, Padova, Pavia, Pisa, Roma 3.

RUSSIA INR Moscow, JINR Dubna.

SPAIN Valencia.

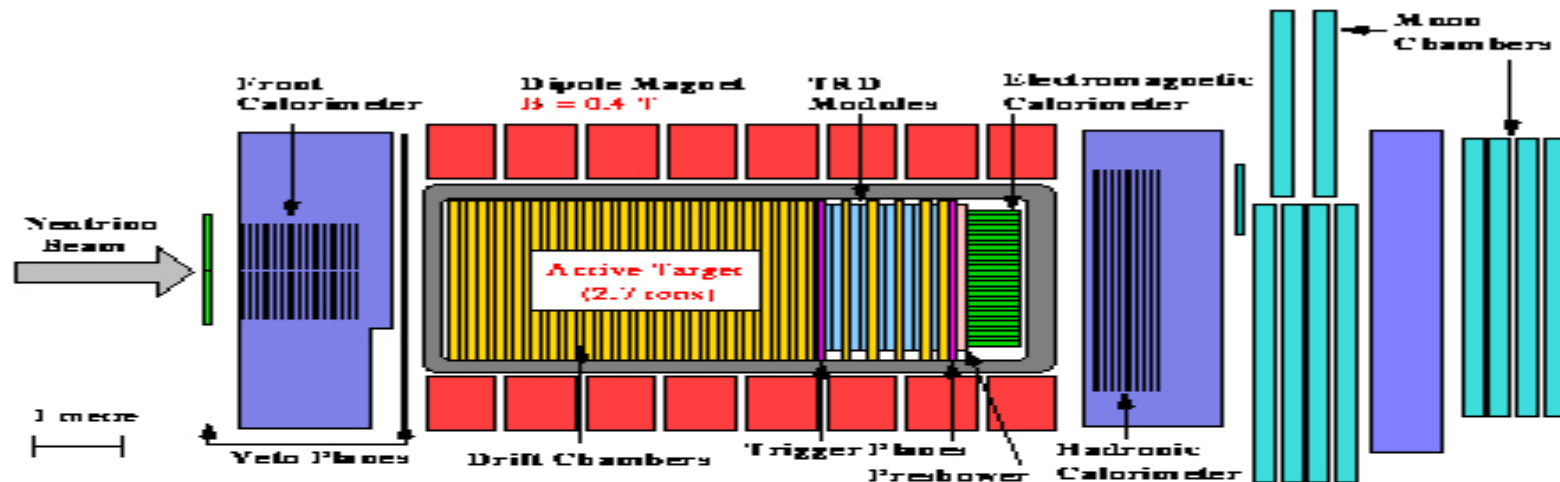
SWITZERLAND Lausanne, ETH Zurich.

USA Harvard, Johns Hopkins, South Carolina, UCLA.

M. Mezzetto Neutrino 2000
P.Astier et al. CERN-EP-2000-049)

NOMAD Detector

(NIM A404(1998)96)



- Drift Chambers (target and momentum measurement)
Spatial resolution $< 200 \mu\text{m}$ (small angle tracks)
Momentum resolution $\sim 3.5\%$ ($p < 10 \text{ GeV/c}$)
- Transition Radiation Detector (TRD) for e^\pm identification
 π rejection $\simeq 10^{-3}$ **$\epsilon(e) \geq 90\%$ for isolated tracks**
- Lead Glass Electromagnetic Calorimeter

$$\frac{\sigma(E)}{E} = 1.0\% + \frac{3.2\%}{\sqrt{E(\text{GeV})}}$$

- Muon Chambers
 $\epsilon \approx 97\%$ for $p_\mu > 5 \text{ GeV/c}$

ν_τ signal extraction technique: excess of events in kinematics box
 \Rightarrow precise energy/momentum & good particle ID

Decay channels

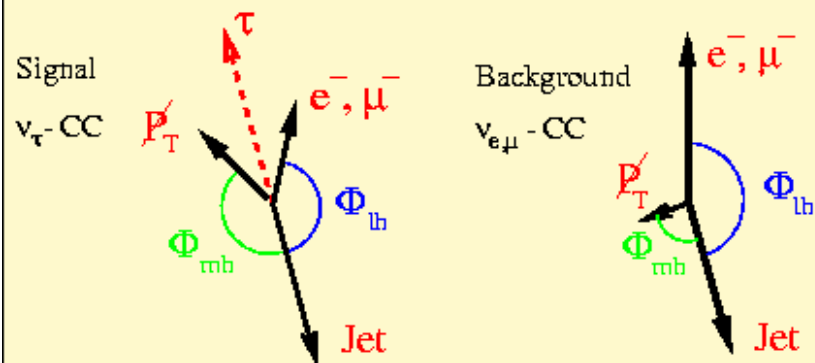
$\tau^- \rightarrow e^- \bar{\nu}_e \nu_\tau$	17.8%	small CC_{ν_e} background
$\tau^- \rightarrow h^- (n\pi^0) \nu_\tau$	49.5%	kinematics very $\neq \text{NC}_\nu$
$\tau^- \rightarrow \pi^+ \pi^- \pi^- (n\pi^0) \nu_\tau$	15.2%	
Total	82.5%	

Not used

$\tau^- \rightarrow \mu^- \bar{\nu}_\mu \nu_\tau$	17.4%
huge CC_{ν_μ} background	

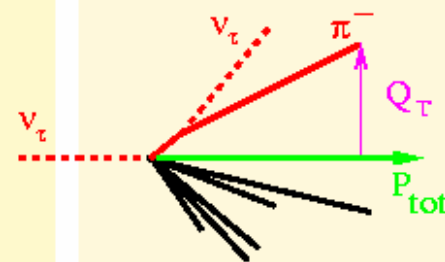
Examples of sensitive kinematics variables

Electron decay, the main source of backgrounds are ν_e CC \Rightarrow kinematics based on the missing momentum and angular relations in the transverse plane.

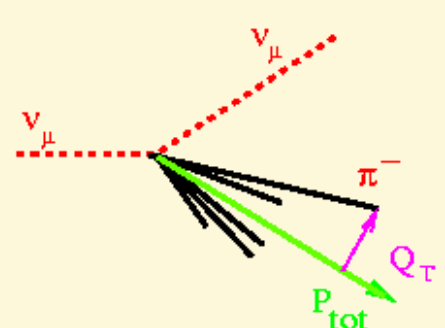


Hadronic decays, the main source of background are neutral current \Rightarrow isolation between the τ visible decay product(s) and the hadronic jet.

Signal



Background from NC



ν_τ signal extraction technique also requires

- Precise simulation of kinematics of signal and background

In kinematics box where signal expected: background known to $\mathcal{O}(10^{-5})$

Data Simulator

- Replace μ^- in CC_{ν_μ} Data (DS) and MonteCarlo (MCS) samples by
Monte-Carlo ν (background NC)

τ^- (signal CC_{ν_τ})

e^- (background CC_{ν_e})

$$\boxed{\mathcal{E}_{Data}^{S,B} = \mathcal{E}_{MC}^{s,b} \frac{\mathcal{E}_{DS}^{S,B}}{\mathcal{E}_{MCS}^{S,B}}}$$

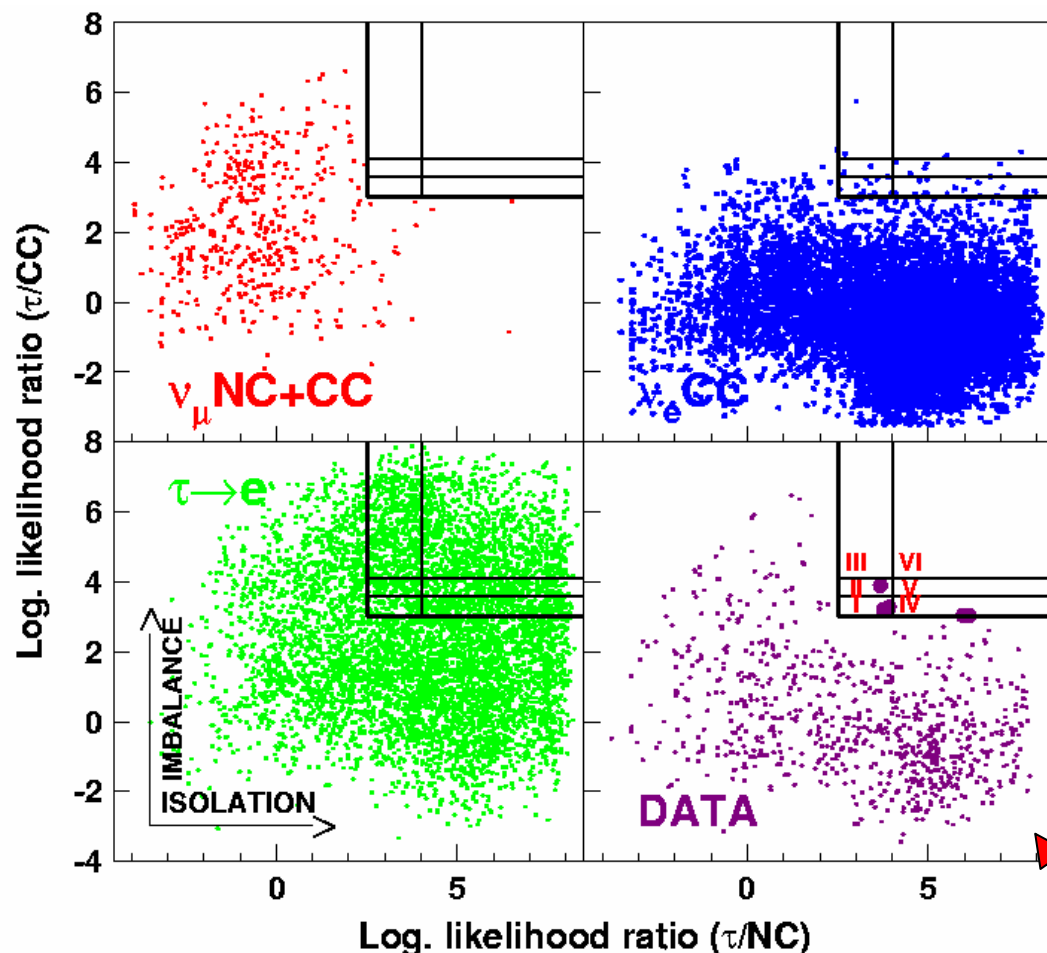
Most systematic (hadron shower simulation, Fermi motion, ...) cancel out

Analysis technique

- **Event classification** (signal/background) **based on log likelihood ratio** $\ln \lambda = \ln \frac{\mathcal{L}_S}{\mathcal{L}_B}$
- $\mathcal{L}_S(X_i)$ ($\mathcal{L}_B(X_i)$) is the probability for an event described by kinematics X_i to be signal (background)
- X_i chosen as most discriminating variables
- \mathcal{L} product of (quasi)independent multivariate $pdf(X_i, X_j, \dots)$
- Tail of $\ln \lambda$ (large signal, small background probs) divided into (independent) signal bins
- **Binning obtained from maximum sensitivity** (based on MC à la Feldman-Cousins)
- **Decay channels and signal bins combined** à la Feldman-Cousins (unified frequentist approach)
- **Blind analysis**: data in potential signal region not looked at until analysis fully defined

$$\tau^- \rightarrow e^- \nu_\tau \bar{\nu}_e$$

Signal region “blindly” selected
divided in 6 bins



Event selection:

efficient e^- ID (20%)

e^-/π^- rejection $\sim 10^6$

e^-/π^0 rejection $\sim 10^4$

Background e^- and γ from NC

$$\mathcal{L}_{\text{NC}} = [[[\theta_{\nu T}, \theta_{\nu H}], \theta_{\min}, Q_T], M_T, E^{\text{TV}}]$$

e^- isolation

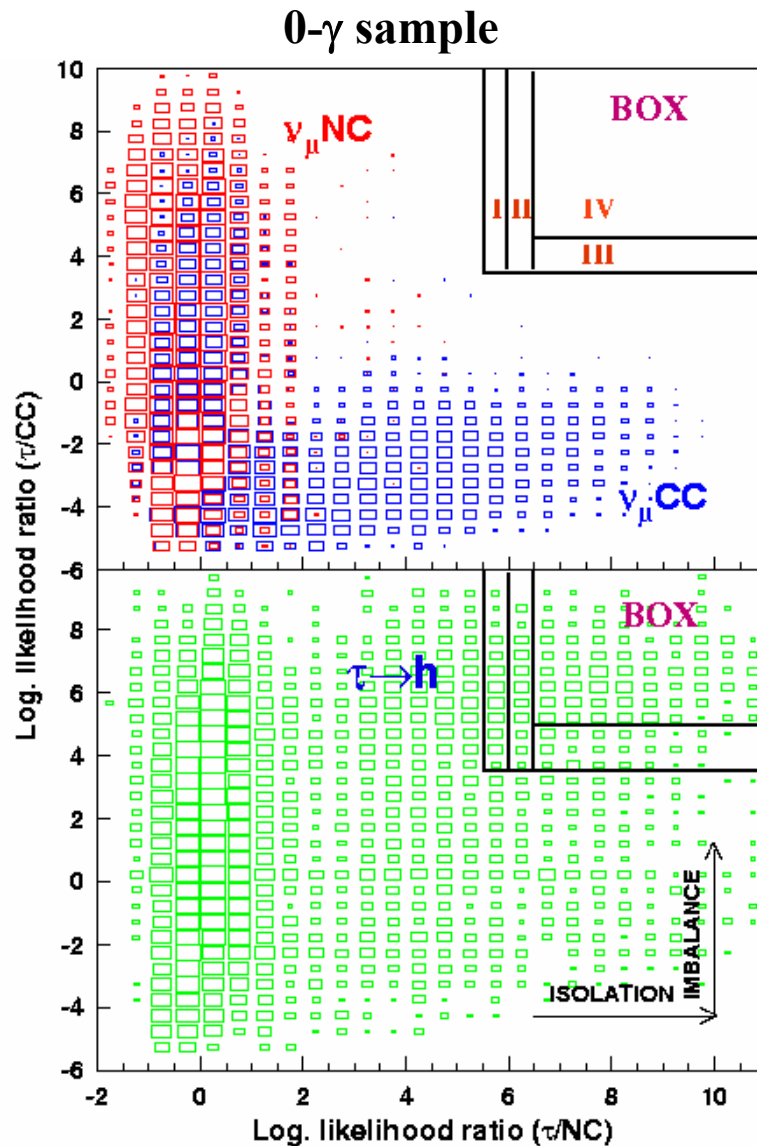
Background e^- from CC_{ν_e}
(1.5% of CC_{ν_μ})

$$\mathcal{L}_{\text{CC}} = [[Q_{\text{Lep}}, \rho_{\tau\nu}, \rho_H], p_T^m, M_T, E_{\text{vis}}]$$

Transverse momentum imbalance

6 events in bins of DATA box
found after box definition

$\tau^- \rightarrow h^- (n\pi^0) \nu_\tau$ inclusive (B.R.=49.5%)



Very new event selection: h^- selection
since CERN-EP 2000-049
shown at Nu2000

π^0 likelihood (2γ)
 ρ likelihood ($>1\gamma$)
 h^- candidate likelihood
better e^-/μ^- rejection

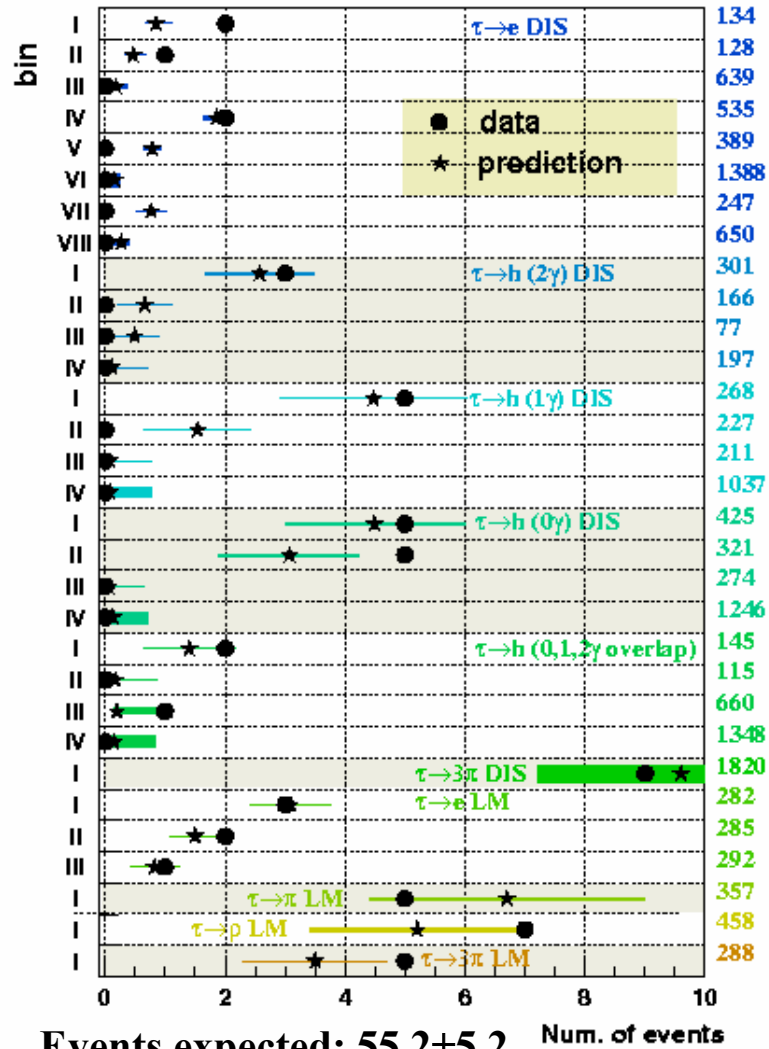
e^- and h^- channels contribute
similarly to sensitivity

$$\mathcal{L}_{NC} = [[[\theta_{\nu T}, \theta_{\nu H}], \theta_{min}, Q_T], p_T^m, p_T^H]$$

$$\mathcal{L}_{CC} = [[I_G, P_T^{lep}/p_T^m, \theta_{\nu, lep}], p_T^m, M_T, E_{vis}]$$

Search Summary

N_τ = expected number of signal events if $P_{osc}(\nu_\mu - \nu_\tau) = 1$



Channels/Bins with very low background

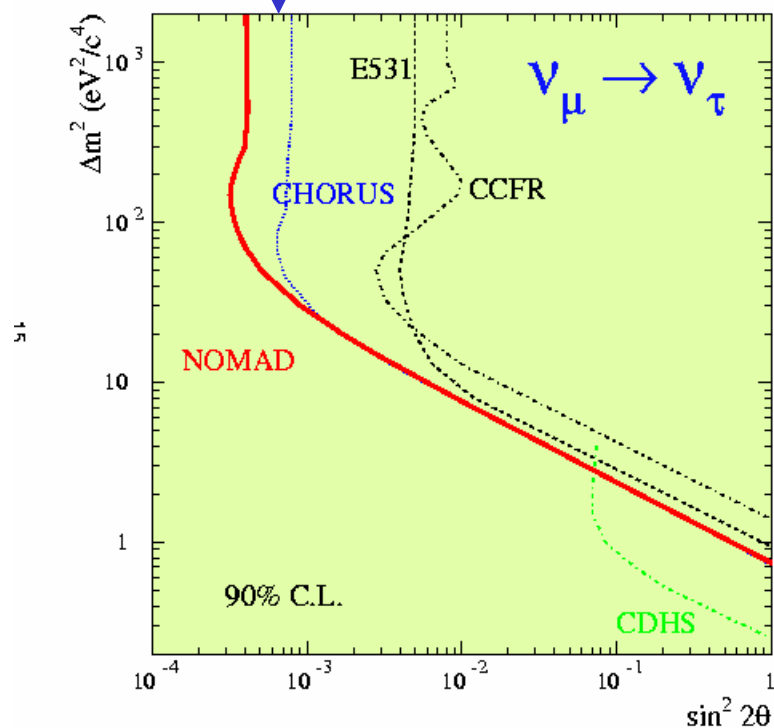
Analysis	Bin	Tot. bkg.	N_τ	Data
$\nu_\tau \bar{\nu}_e e$	DIS III	$0.28^{+0.31}_{-0.09}$	903	0
$\nu_\tau \bar{\nu}_e e$	DIS VI	0.25 ± 0.09	1694	0
$\nu_\tau h(0\gamma)$	DIS III	$0.05^{+0.60}_{-0.03}$	274	0
$\nu_\tau h(0\gamma)$	DIS IV	$0.12^{+0.60}_{-0.05}$	1246	0
$\nu_\tau h(1\gamma)$	DIS III	$0.07^{+0.70}_{-0.04}$	211	0
$\nu_\tau h(1\gamma)$	DIS IV	$0.07^{+0.70}_{-0.04}$	1037	0
$\nu_\tau h(2\gamma)$	DIS IV	$0.11^{+0.60}_{-0.06}$	197	0
$\nu_\tau h(\text{overl.})$	DIS III	$0.20^{+0.70}_{-0.06}$	660	1
$\nu_\tau h(\text{overl.})$	DIS IV	$0.14^{+0.70}_{-0.06}$	1348	0
$1.29^{+1.60}_{-0.18}$				1

NOMAD exclusion plots (preliminary)

@90% C.L.
Unified frequentist
approach

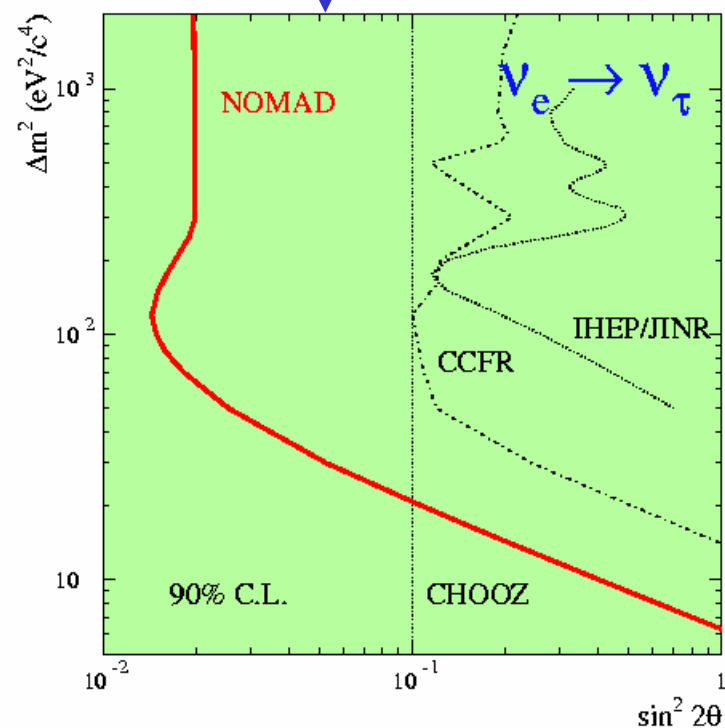
Still room for progress, i.e. $\tau \rightarrow \nu_\tau$ 3π DIS analysis.

New CHORUS (different statistical method)



$$P(\nu_\mu \rightarrow \nu_\tau) \leq 2.03 \cdot 10^{-4}, \text{ sensitivity : } 2.6 \cdot 10^{-4}$$

$$\sin^2 2\theta_{\mu\tau} > 4.06 \cdot 10^{-4} \quad \text{for large } \Delta m^2$$



Computed with the proper ν_e flux, 100 times smaller than the ν_μ flux.

$$P(\nu_e \rightarrow \nu_\tau) \leq 1.00 \cdot 10^{-2}, \text{ sensitivity: } 1.3 \cdot 10^{-2}$$

More to expect from NOMAD:

$\nu_\mu \rightarrow \nu_\tau$ and $\nu_e \rightarrow \nu_\tau$ still being improved

$\tau^- \rightarrow \pi^- \pi^- \pi^+ \nu_\tau$ channel

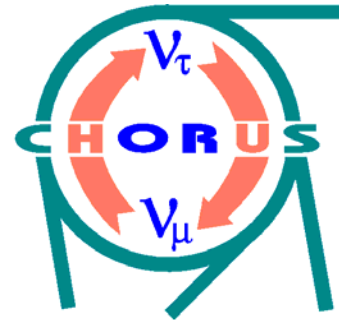
$\nu_\mu \rightarrow \nu_e$ in progress

talk by R.Petti in PS6

Λ polarization in $\nu_\mu N \rightarrow \mu^- \Lambda X$

talk by Minh-Tam Tran in PS6

CHORUS experiment
Search for $\nu_\mu \rightarrow \nu_\tau$ oscillation



Collaboration

**Belgium (Brussels, Louvain-la-Neuve),
CERN,
Germany (Berlin, Münster),
Israel (Haifa),
Italy (Bari, Cagliari, Ferrara, Naples, Rome, Salerno),
Japan (Toho, Kinki, Aichi, Kobe, Nagoya, Osaka, Utsunomiya) ,
Korea (Gyeongsang),
The Netherlands (Amsterdam),
Russia (Moscow),
Turkey (Adana, Ankara, Istanbul)**

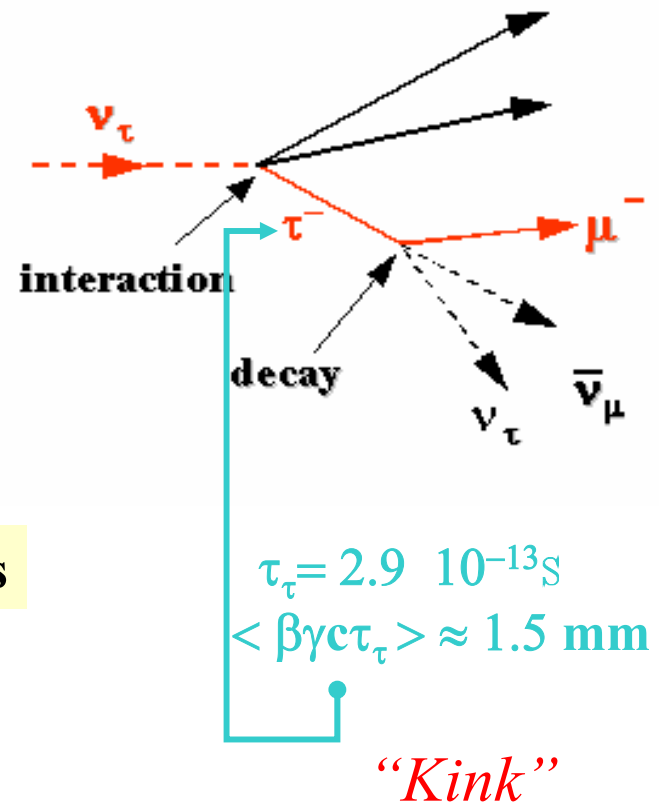
**L.Ludovicci Neutrino 2000
E.Eskut at al. CERN-EP-2000-0??)**

ν_τ Direct detection technique

Observation of the τ -lepton track produced in CC ν_τ interactions in 770 kg nuclear emulsion target : **“kink” topology**

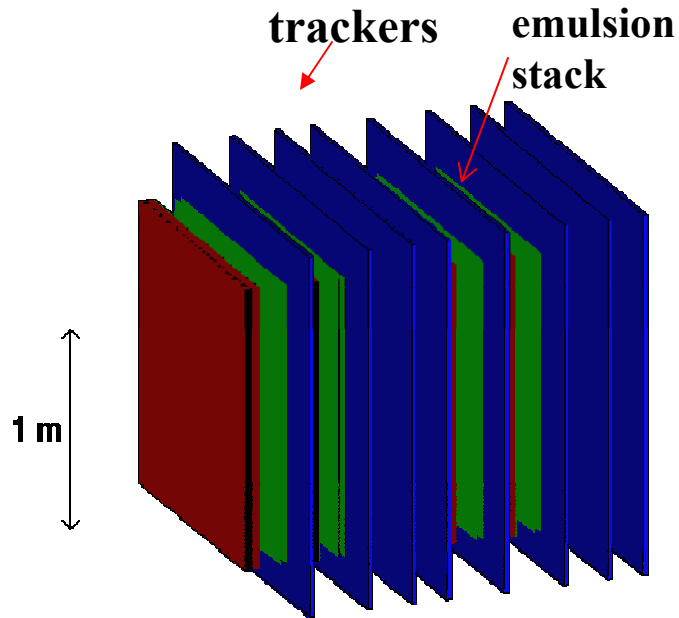
$\mu^- \nu_\tau \bar{\nu}_\mu$	B.R.	18%
$h^- \nu_\tau (n \pi^0)$		50%
$\pi^- \pi^- \pi^+ \nu_\tau (n \pi^0)$	14%	not yet included

Prompt ν_τ background $< \sim 0.1$ event in 4 years

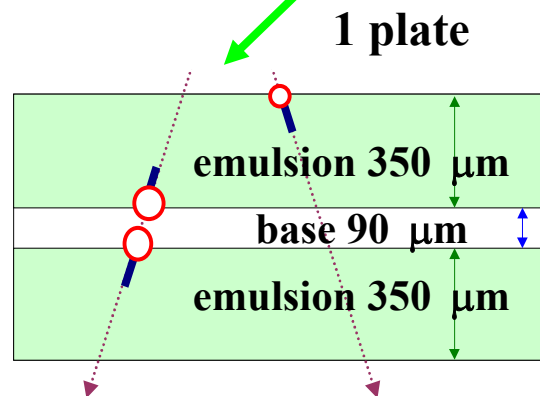


CHORUS target

4 stacks of emulsion interleaved with
fibre trackers and
emulsion interface trackers

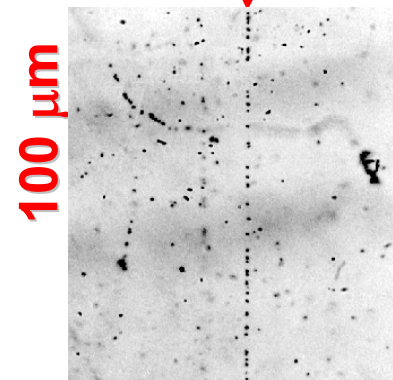


- 1 stack : $142 \times 144 \times 2.8 \text{ cm}^3$
= 36 plates of $790 \text{ }\mu\text{m}$



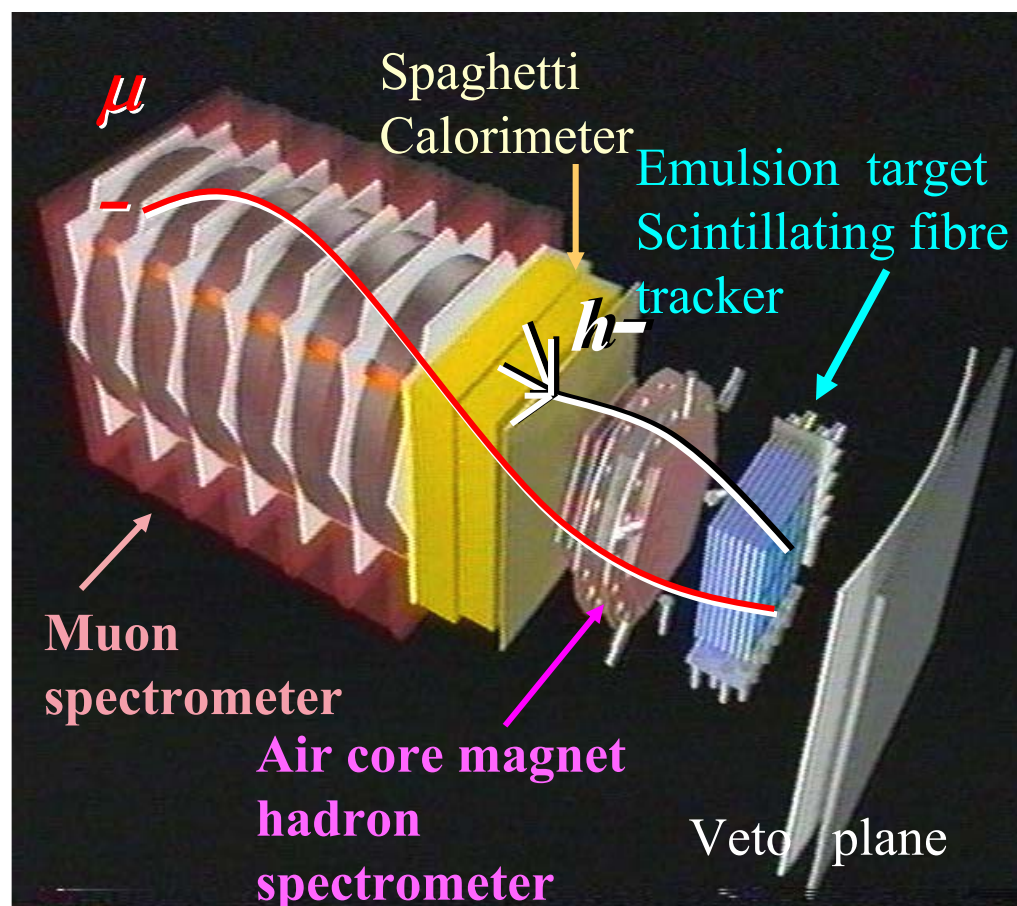
Grain size $\sim 1. \text{ }\mu\text{m}$
Grain measurement $\sim 0.3 \text{ }\mu\text{m}$
Angular resolution $\sim 1.5 \text{ mrad}$

MIP
 $\sim 35 \text{ grains} / 100 \text{ }\mu\text{m}$



80 μm

CHORUS detector : event kinematics measurement



Hadron Sign and momentum

Air-core magnet hadron spectrometer

$$\Delta p/p = \sqrt{(0.035 \cdot p(\text{GeV}/c) + 0.22^2)}$$

Showers energy, missing P_t

Lead&fibers "spaghetti" calorimeter

$$\Delta E/E = 32\%/\sqrt{E} \text{ (hadrons)}$$

$$\Delta E/E = 14\%/\sqrt{E} \text{ (electrons)}$$

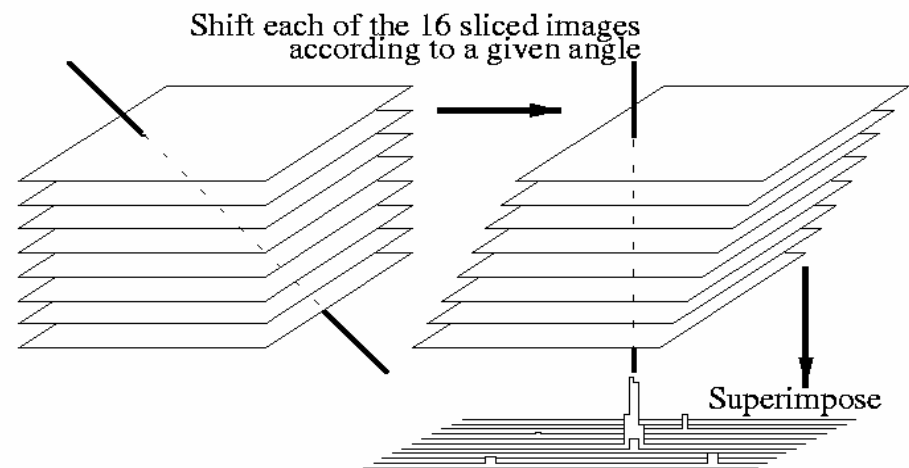
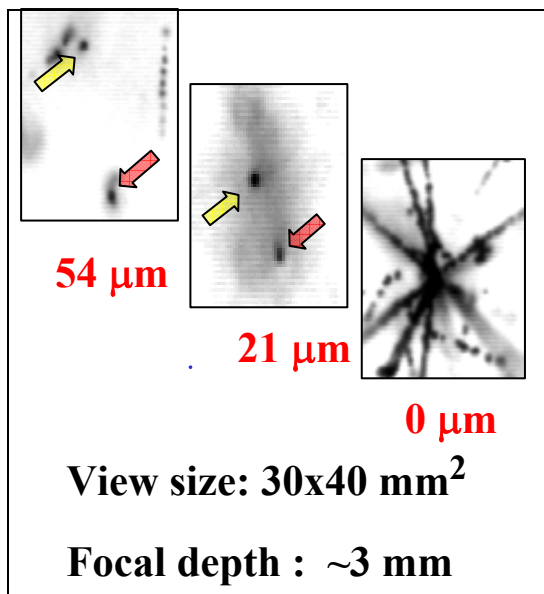
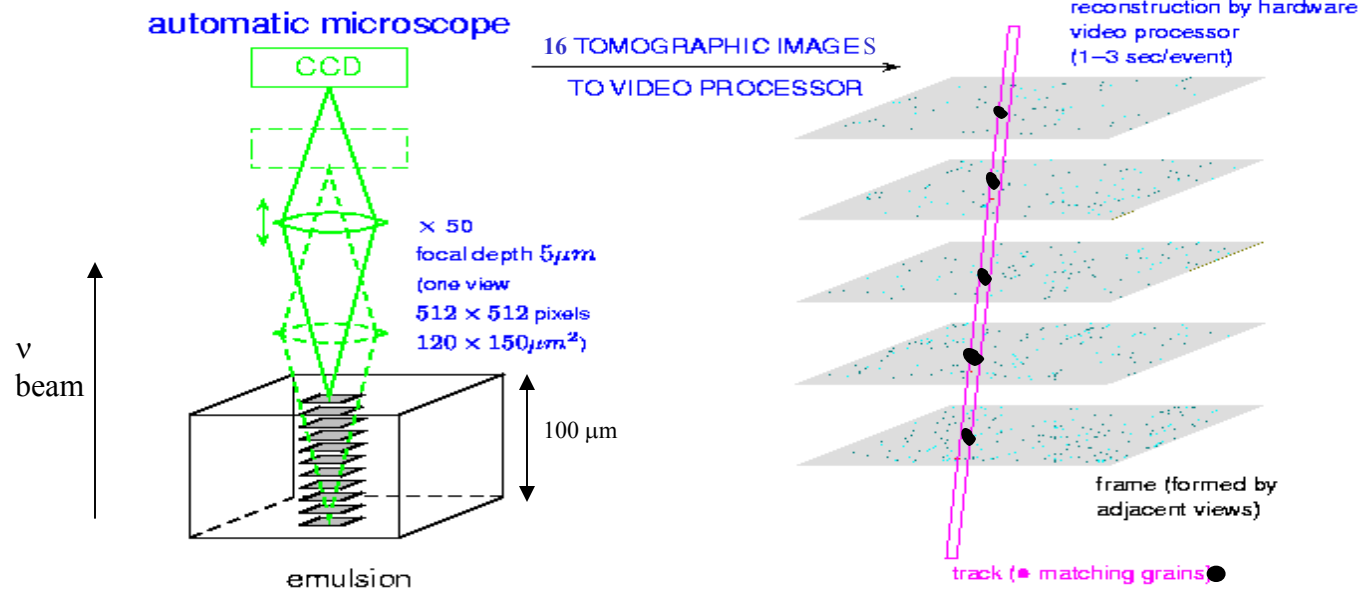
$$\Delta q_{\text{hadr}} \sim 60 \text{ mrad @ } 10 \text{ GeV}$$

Muon ID, sign and momentum

Iron-core muon spectrometer

$$\Delta p/p \sim 10\%-15\% \text{ (} p < 70 \text{ GeV)}$$

Automatic Emulsion Data Taking (K.Niwa and Nagoya University)



Analysis strategy

- Event reconstruction and loose kinematics selection

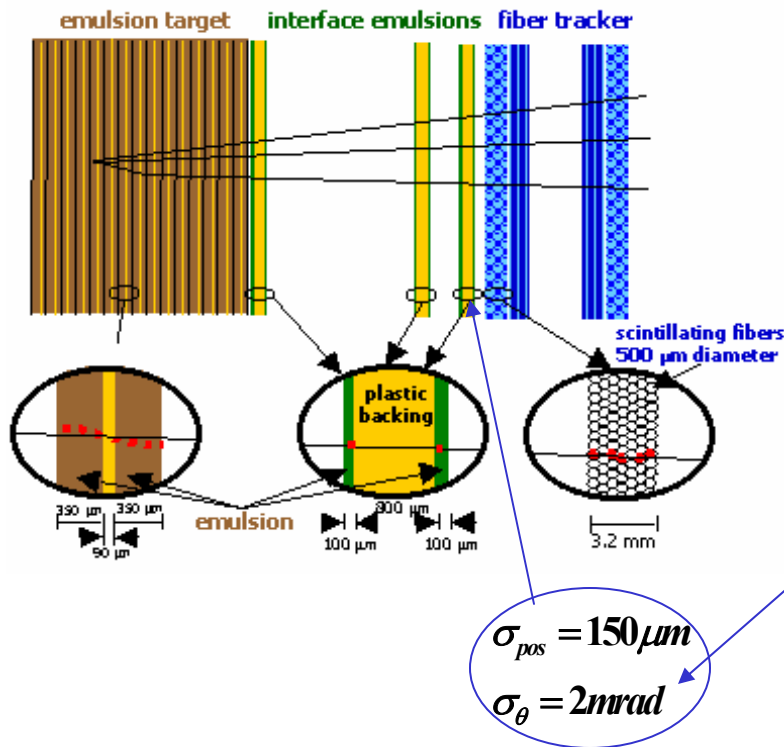
$$\mu^- \nu_\tau \bar{\nu}_\mu$$

1 μ^- with $p_\mu < 30 \text{ GeV}/c$

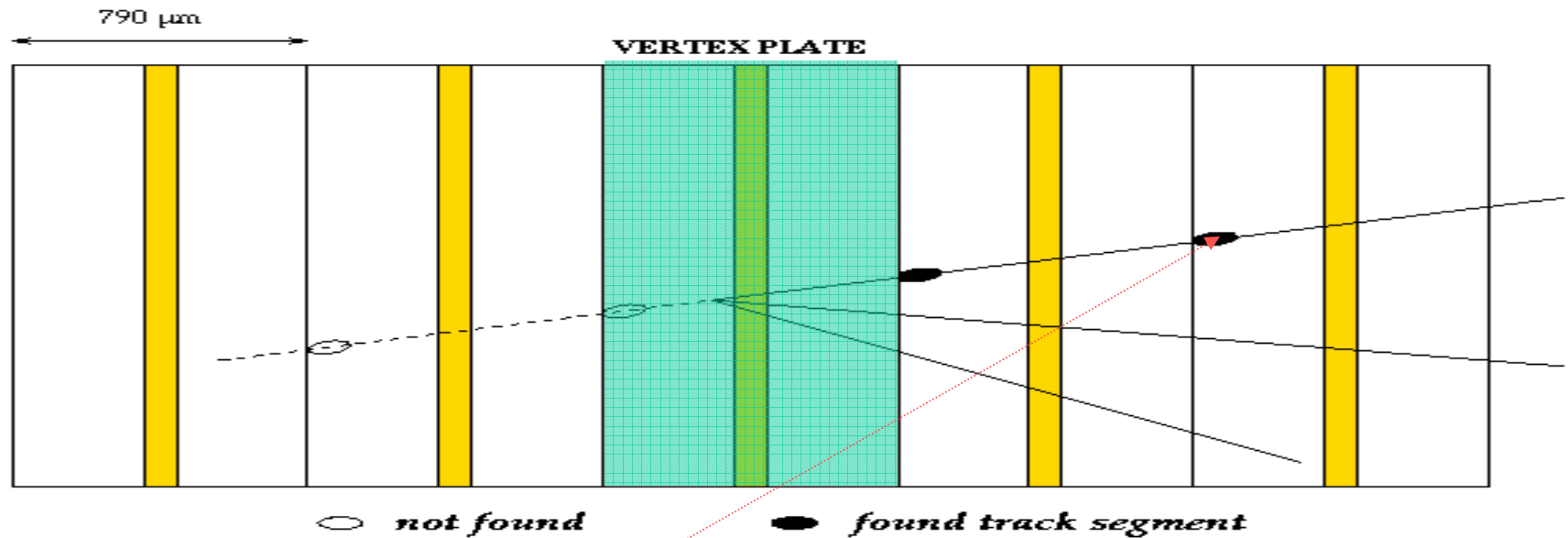
$$h^- \nu_\tau (n \pi^0)$$

no μ^- and at least 1 h^- with $1 < p < 20 \text{ GeV}/c$

- Track predictions at emulsion trackers for tracks reconstructed in scintillating fibre trackers
- Tracks found and followed by automatic microscopes in 3 successive interface emulsion trackers up to stack entry
- Followed back plate by plate in target to find vertex
- Automatic search for a “kink” decay topology: 3% of events
- Events with “kink” are analysed manually: 1% of selected events retained as candidates
- Precise kinematics analysis of candidates

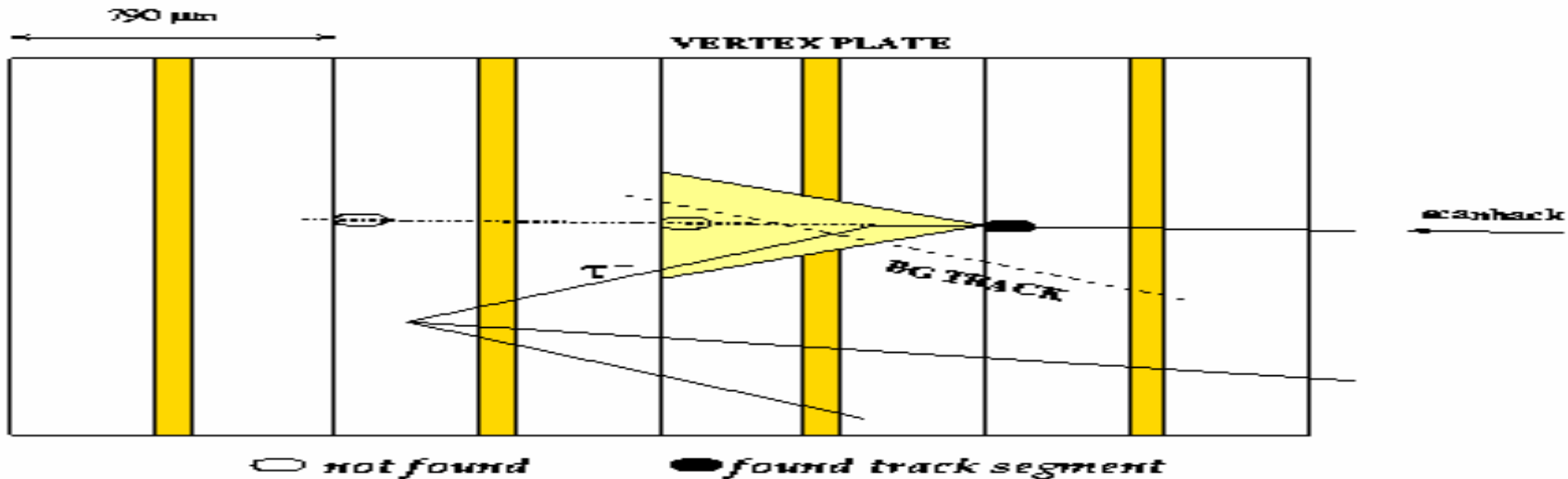


Automatic Vertex Location



- Follow-up track, plate by plate to the vertex
- **100 μm** most upstream of each target plate are scanned
- Vertex defined by the first plate out of two consecutive plates where a track segment is not found

Kink Finding - Parent Search (Large Angle-Long Path kinks)



$100\ \mu\text{m}$ most upstream of the vertex plate are searched for all track segments in a cone of width $\propto 1/P$

Segments with small impact parameters w.r.t. the follow-up track
→ **Candidate track parent track**
→ **Manual microscope inspection (3% of scanned events)**

Manual candidate event selection:

• “clean” 1-prong kink: no sign of nuclear interaction



1% of inspected events

Data

<i>Year</i>	<i>1994</i>	<i>1995</i>	<i>1996</i>	<i>1997</i>	<i>All</i>
POT / 10^{19}	0.81	1.20	1.38	1.67	5.06
Good emulsion	97%	73%	100%	100%	~93%
Expected Ncc / 10^3	120	200	230	290	840
Emulsion trigger / 10^3	422	547	617	719	2,305
1μ to be scanned	66,911	110,916	139,527	151,105	468,459
1μ scanned so far	88%	55%	81%	83%	77%
1μ vertex location and kink search	20,400	21,610	41,558	52,789	136,357
0μ to be scanned	19,846	29,350	37,143	36,073	122,412
0μ scanned so far	60%	58%	79%*	67%*	67%
0μ vertex location and kink search	3,024	4,424	8,704	7,054	23,206



Background evaluation and reduction

- π and K decays

- P_t (daughter-parent) > 250 MeV/c : reject 100 %

- Charm background

- primary lepton not identified and, if D^+ , charge of secondary wrong
- D^- produced by ν_μ/ν_e beam component

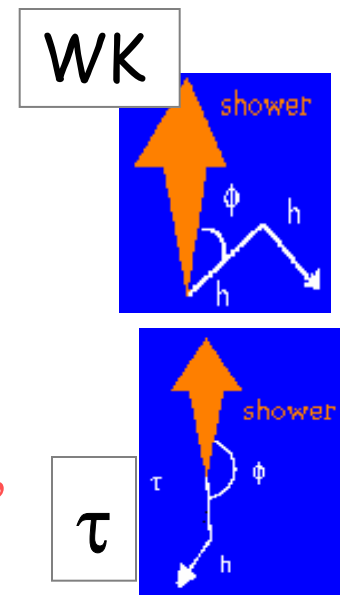
- “White kink

- hadron elastic scattering with no sign of nuclear activity
- badly known rate is **measured** at large distance from vertex
- number within τ^- decay path computed by MC

- Charm decays and white kinks reduction in the h- channel

- P_h dependent τ candidate decay path cut
such that 80% of the true τ are retained $\forall P_h$
- $\Phi(\tau\text{-Hadron shower})$ in transverse plane > 90 °

Cuts optimised for maximum sensitivity “without looking at data”



Background	1μ	0μ
D^- from $\bar{\nu}_e / \bar{\nu}_\mu$ CC with primary μ^+ / e^+ missed	0.11	0.03
D^+ from $\bar{\nu}_e / \bar{\nu}_\mu$ CC with primary μ^- / e^- missed and wrong charge for decay μ^+ / h^+	0.03	0.30
CC and NC associated D^+ / D^0 missed and $D^- \rightarrow \mu^- / h^- + \text{neutrals}$	negligible	
"White kinks" h^- elastic scattering with no nuclear activity	0	0.8
Prompt beam ν_τ from D_S decays	0.05	0.05

Results

$$P_{\mu\tau}^{osc} = N_{\tau} / N_{\tau}^{max}$$

$$N_{\tau}^{max} (P_{\mu\tau}^{osc} = 1) = \sum_{i=\mu^-, h^-} BR_i \cdot N_i < \frac{\sigma_{\nu\tau}^{cc}}{\sigma_{\nu\mu}^{cc}} \frac{A_{\tau i}}{A_{\mu i}} \varepsilon_{\tau i} > = 5014 + 2004 = 7018$$

branching
ratios

Nb
events

cross--section
ratios

location efficiency
ratios

kink finding
efficiency

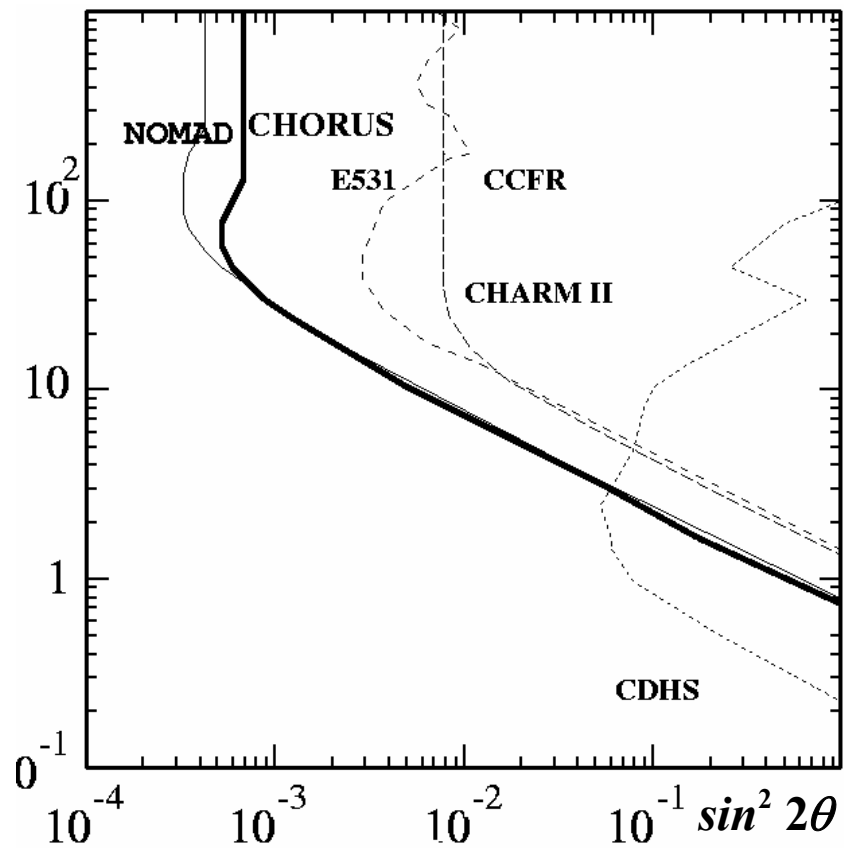
Channel	Observed	Expected background	N_{τ}^{max}
μ^-	0	0.1	5 014
h^-	0	1.1	5 004

Upper limit on N_t @ 90% C.L. = 2.4
(T.Junk, NIM A434 (1999) 435)



$$P_{\mu\tau}^{osc} < 3.4 \cdot 10^{-4}$$

$$\nu_{\mu} \rightarrow \nu_{\tau}$$

 $\Delta m^2 (eV^2)$


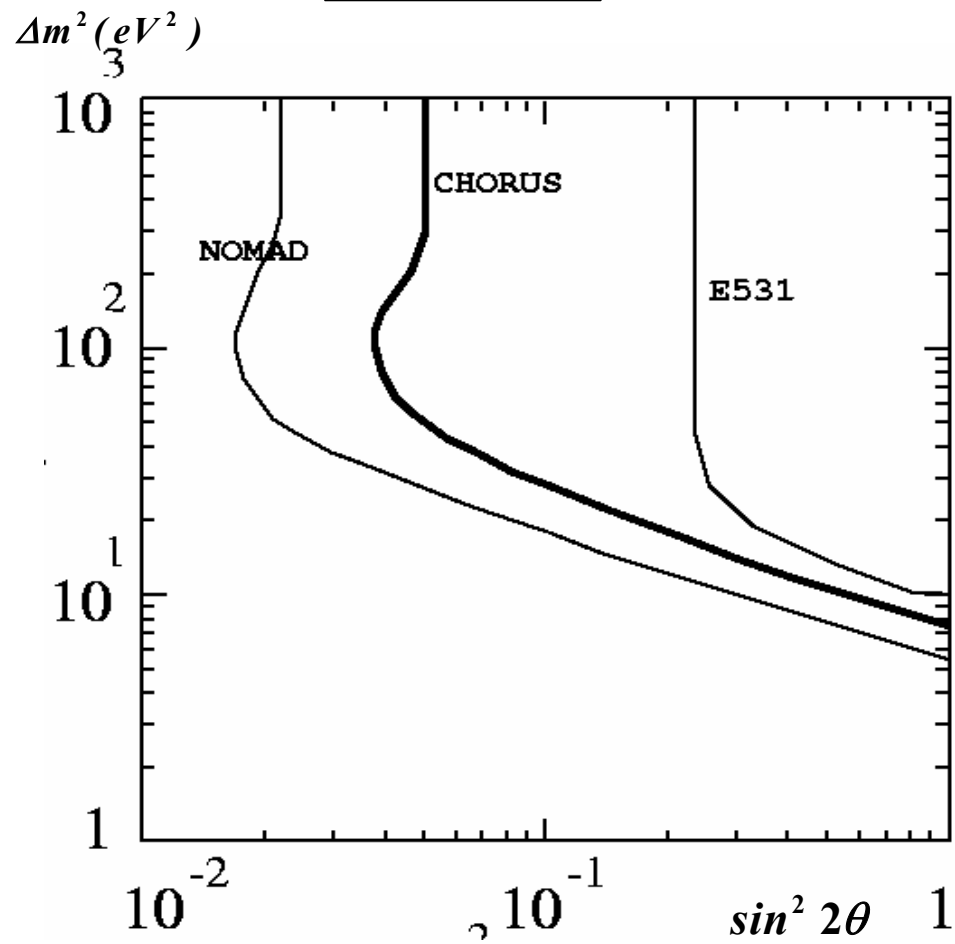
Junk $N_{\tau} < 2.4$
 Feldman_Cousins $N_{\tau} < 1.4$

} @ 90% C.L. given { 1.2 expected background
 0 observed

	NOMAD (Feldman-Cousins)	CHORUS (Junk)	CHORUS (F.-C.) "to compare"
sensitivity (F.C.)	$2.6 \cdot 10^{-4}$	$3.7 \cdot 10^{-4}$	$3.7 \cdot 10^{-4}$
$P_{\nu\tau}^{osc}$	$2.03 \cdot 10^{-4}$	$3.4 \cdot 10^{-4}$	$2.0 \cdot 10^{-4}$
$\sin^2 2\theta$ @ large Δm^2	$4.06 \cdot 10^{-4}$	$6.8 \cdot 10^{-4}$	$4.0 \cdot 10^{-4}$
$\Delta m^2 (eV^2)$ @ full mixing	0.6	0.6	0.6

see talk by R.Petti in PS6

$$\nu_e \rightarrow \nu_\tau$$



There are 0.9% of events in the beam

$$P_{e\tau}^{osc} < 2.6 \cdot 10^{-4}$$

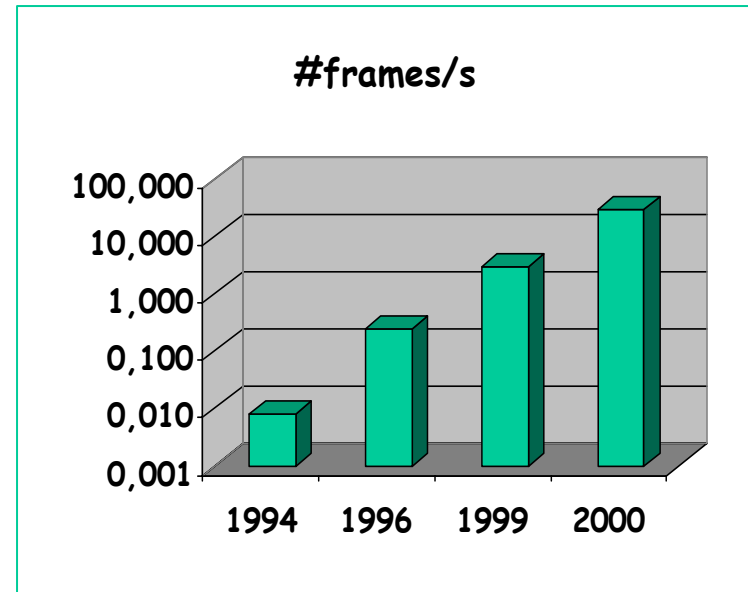
**More to expect from CHORUS:
2 years Phase-2 analysis launched**



Reach $P_{\text{osc}} < 10^{-4}$

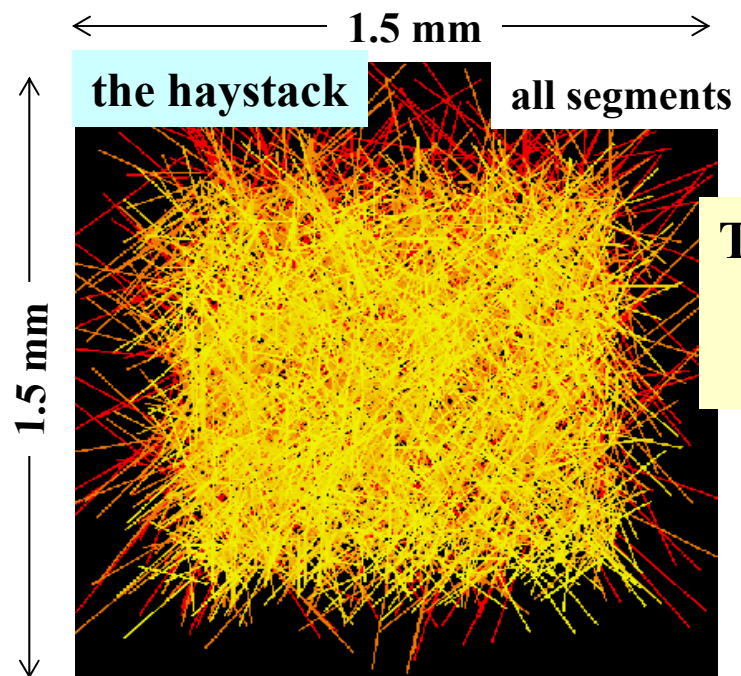
Among other things:

- **improved kink (τ) finding efficiency**
 - **3-hadron and e^- decay channels**
- in emulsion thanks to new upgrade
in automatic microscope technology



Other Physics

- CHARM physics in emulsion (D_s^{*+} observation published)
and in calorimeter as target (J/ψ production submitted)
- Form factors
- Trident ($\mu^+ \mu^- \nu$) production
- Search for heavy neutral leptons ...



The Net Scan

Real Data

